

# Constraints on Majorana Dark Matter from LHC and IceCube

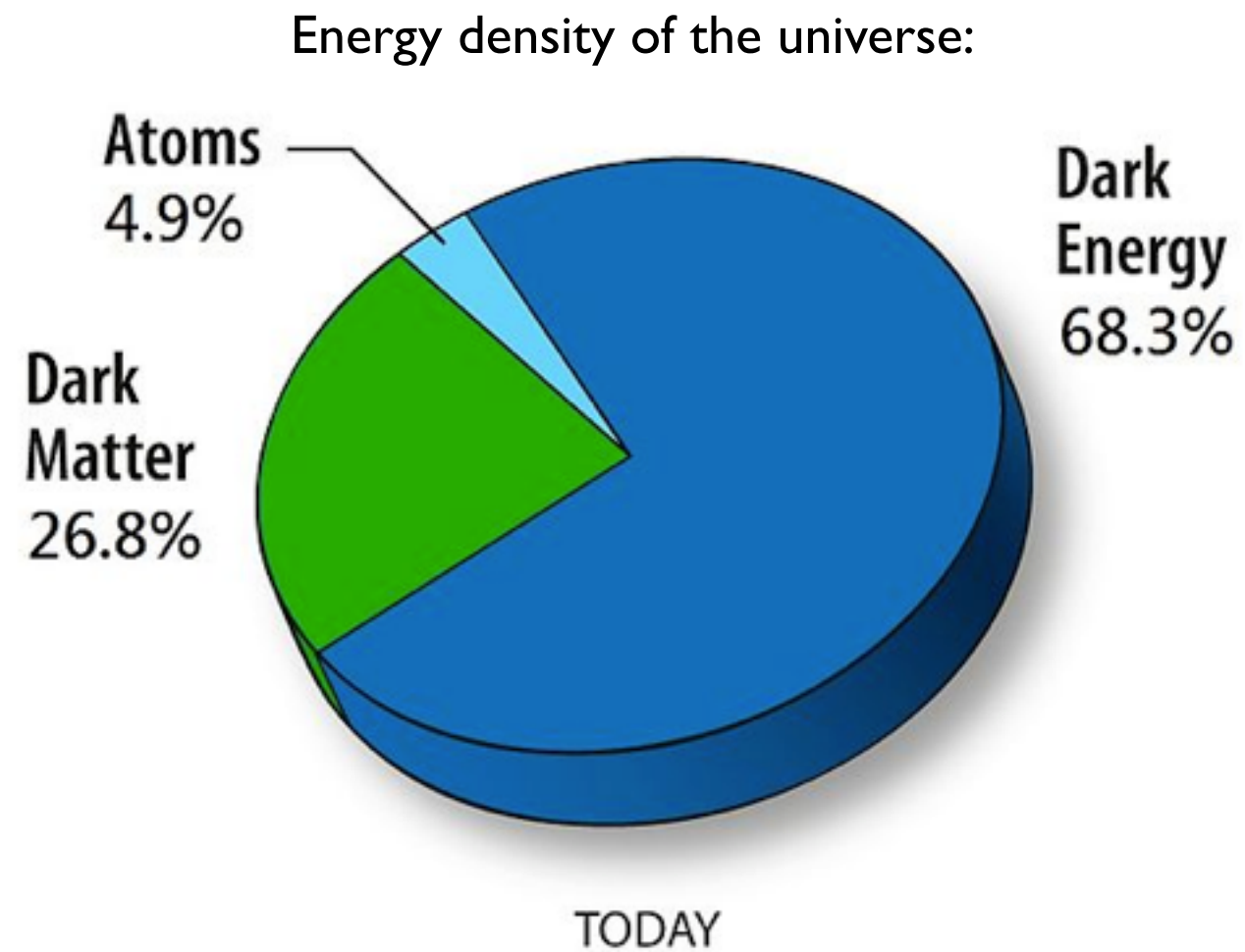
[based on arXiv:1509.07867; JH, Michael Krämer, Mathieu Pellen, Christopher Wiebusch]

Jan Heisig (RWTH Aachen University)



Meeting of Research Unit  
**New Physics at the LHC**  
Bonn, October 28<sup>th</sup>, 2015

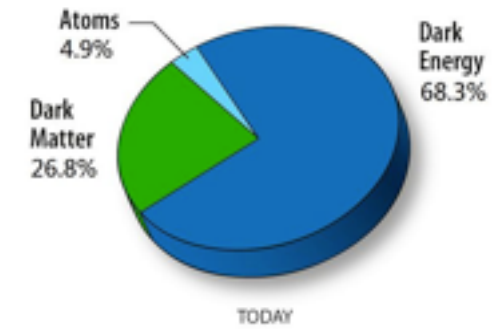
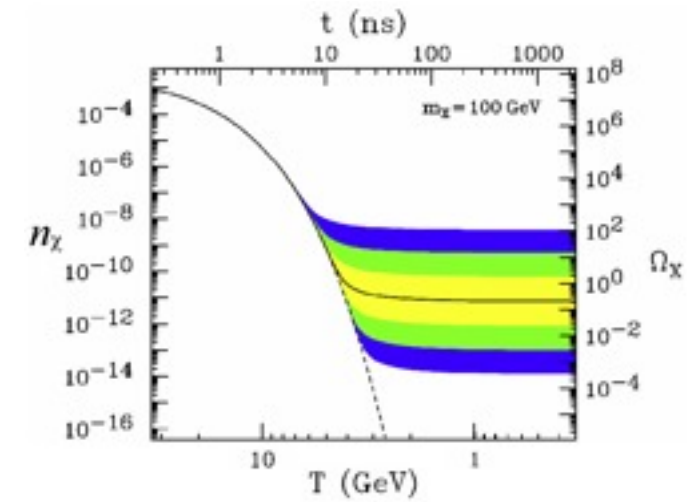
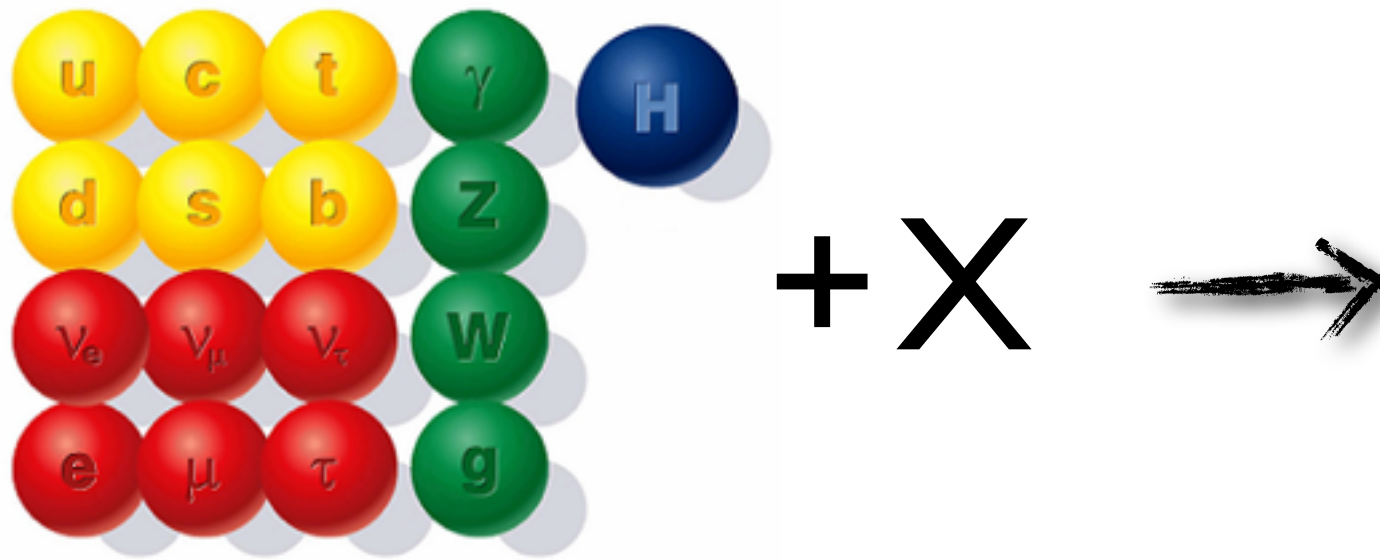
# Dark Matter → ?



# Dark Matter → WIMP

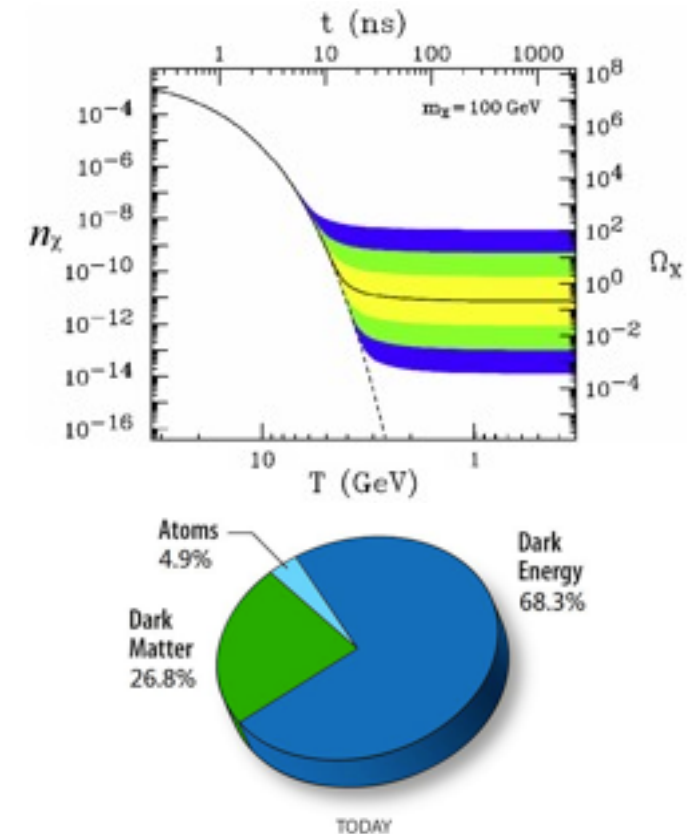
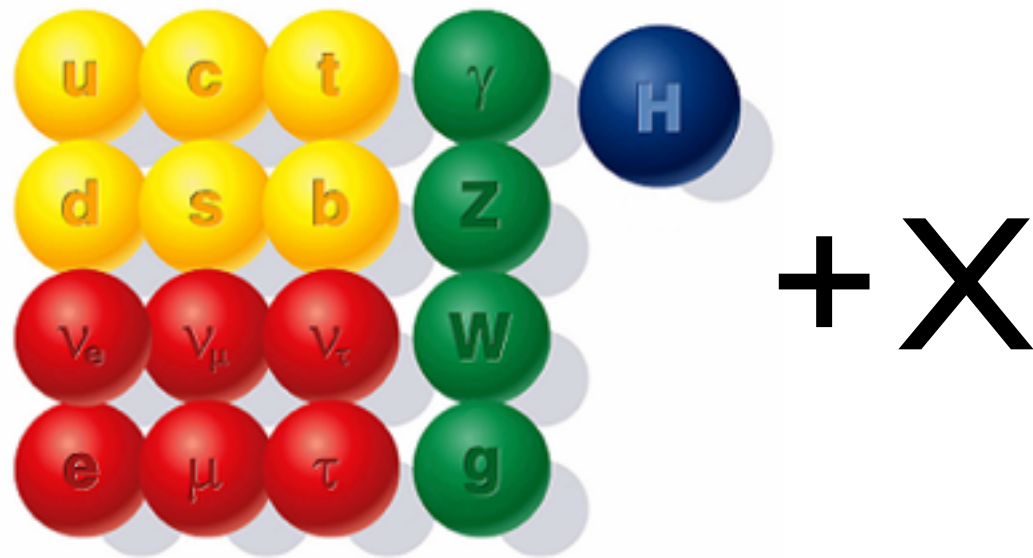


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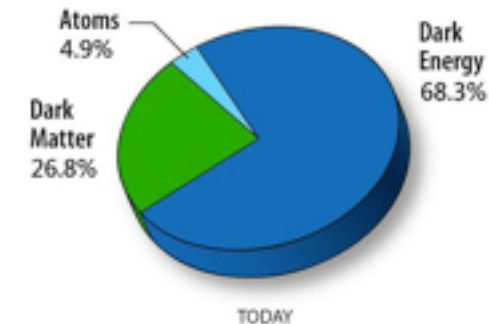
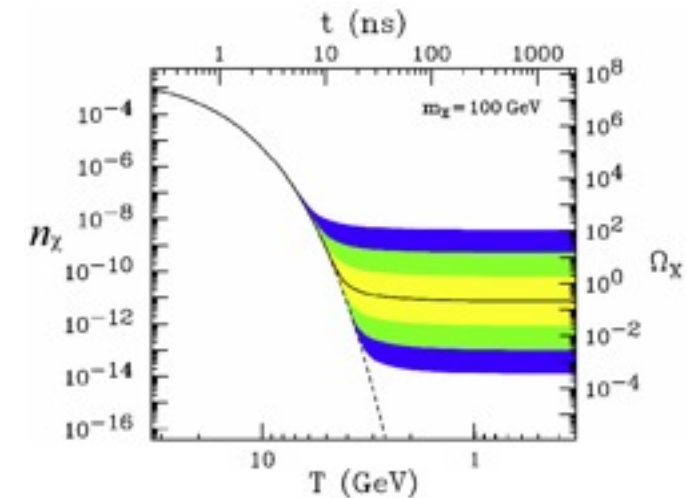
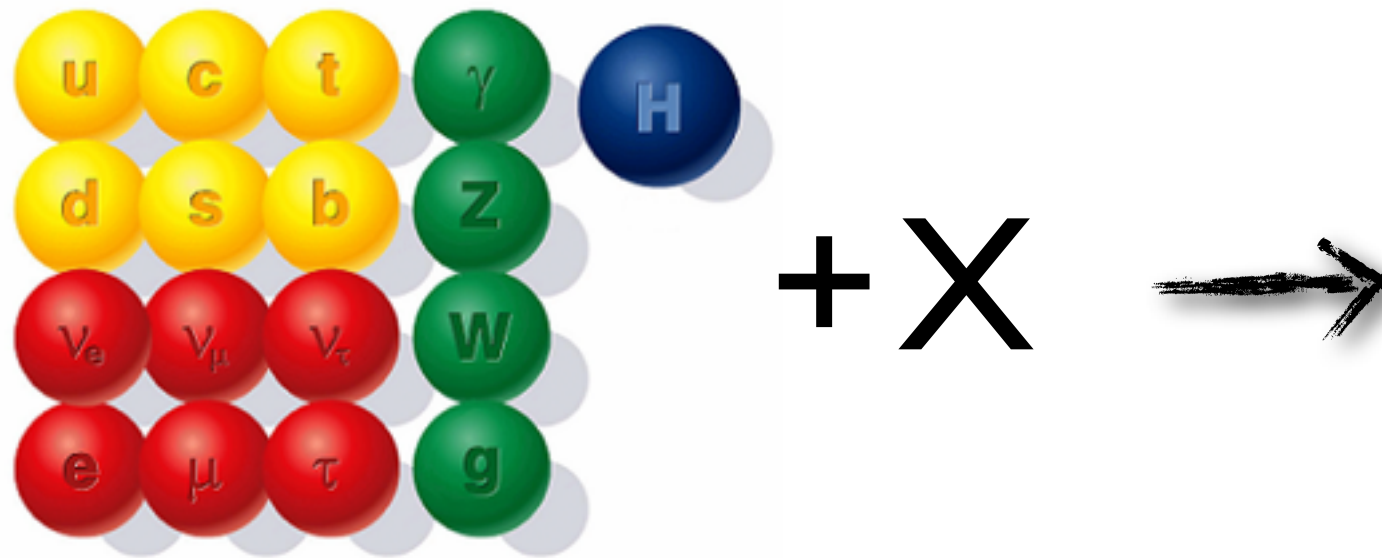
## Pheno description

- Effective Operators
- Simplified Models

## Probe

- Direct detection experiments
- Indirect detection (IceCube)
- Thermal relic density
- DM production@LHC

# Dark Matter → WIMP



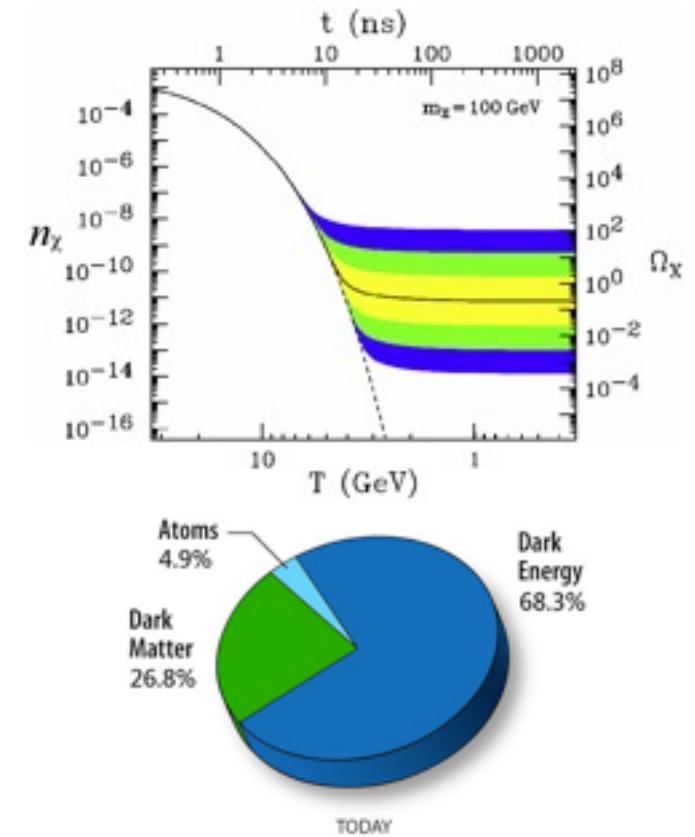
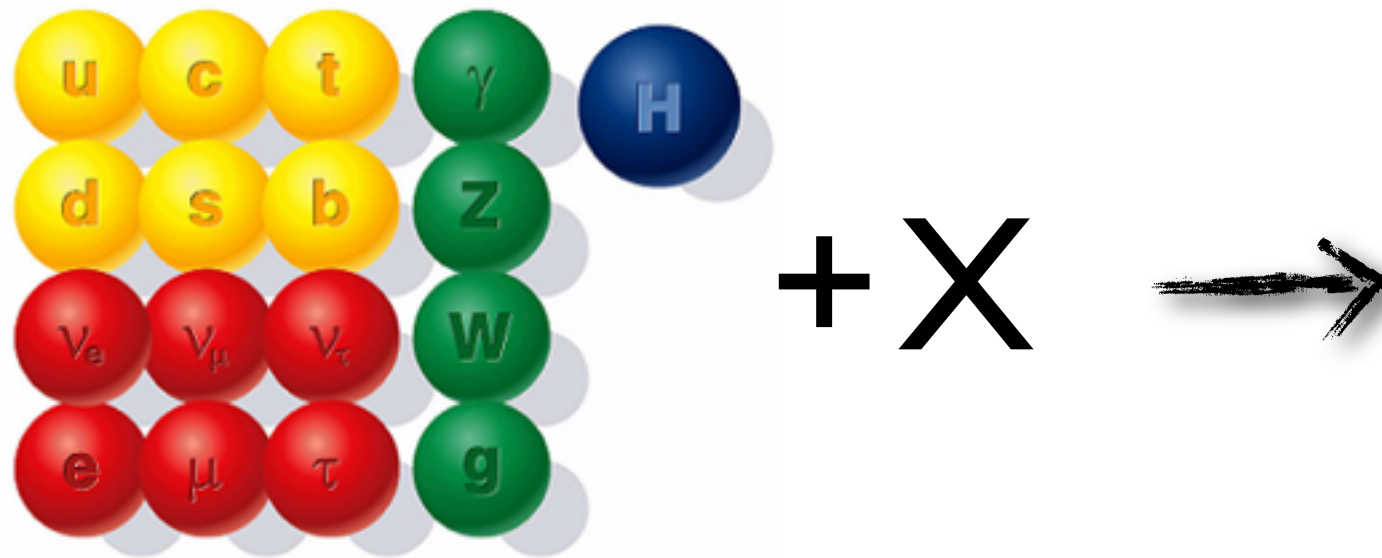
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# Simplified Models for Dark Matter

Busoni, De Simone, Morgante, Riotto: 1307.2253

Buchmueller, Dolan, McCabe: 1308.6799

Busoni, De Simone, Jacques, Morgante, Riotto: 1405.3101

Buchmueller, Dolan, Malik, McCabe: 1407.8257

Harris, Khoze, Spannowsky, Williams: 1411.0535

Chala, Kahlhoefer, McCullough, Nardini, Schmidt-Hoberg: 1503.05916

Backović, Krämer, Maltoni, Martini, Mawatari, Pellen: 1508.05327

Baker, Brod, Hedri, Kaminska, Kopp, Liu, Thamm, Vries, Wang, Yu, Zurita: 1510.03434

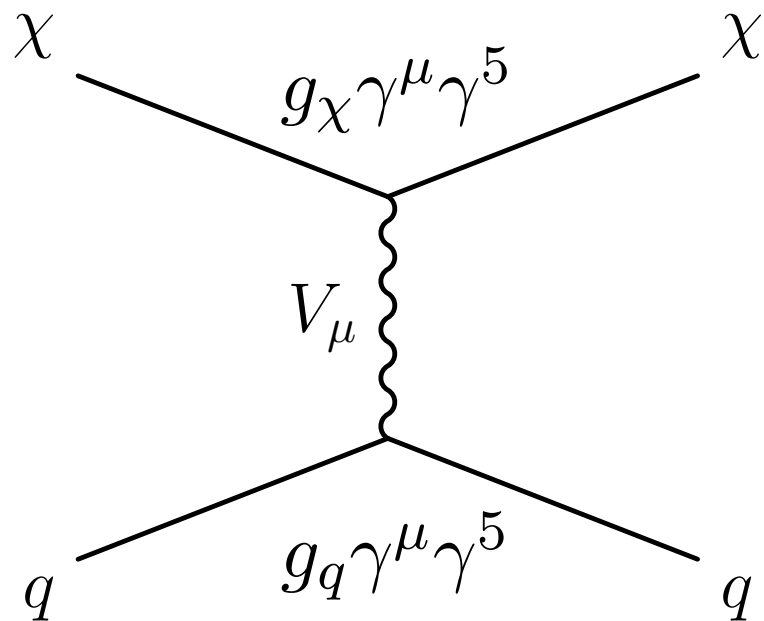
...

## This Talk

- Explore complementarity new Limits from LHC and IceCube
- Discuss EFT ↔ Simplified Model

# A "direct-detection-phobic" model

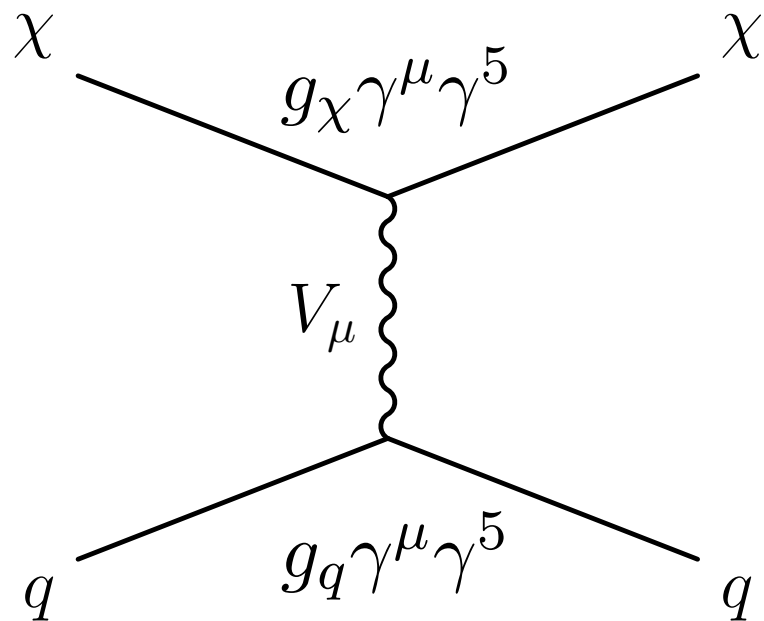
- Model where LHC and IceCube are competitive
  - No spin-independent WIMP-nucleon scattering



- Vector-boson s-channel messenger
- Majorana DM
- Axial couplings to quarks and DM (no couplings to leptons)
- Four parameters:  $M_V, m_\chi, g_q, g_\chi$

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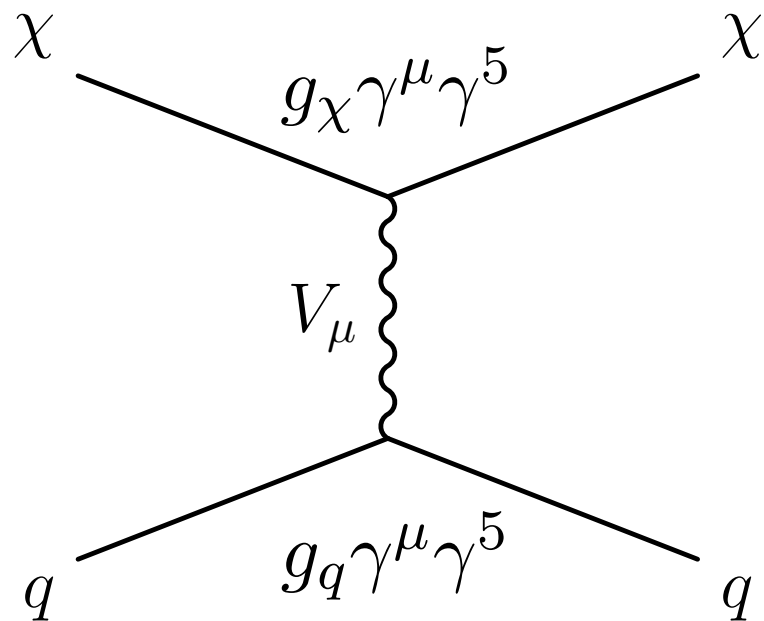
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Note: not all values allowed!

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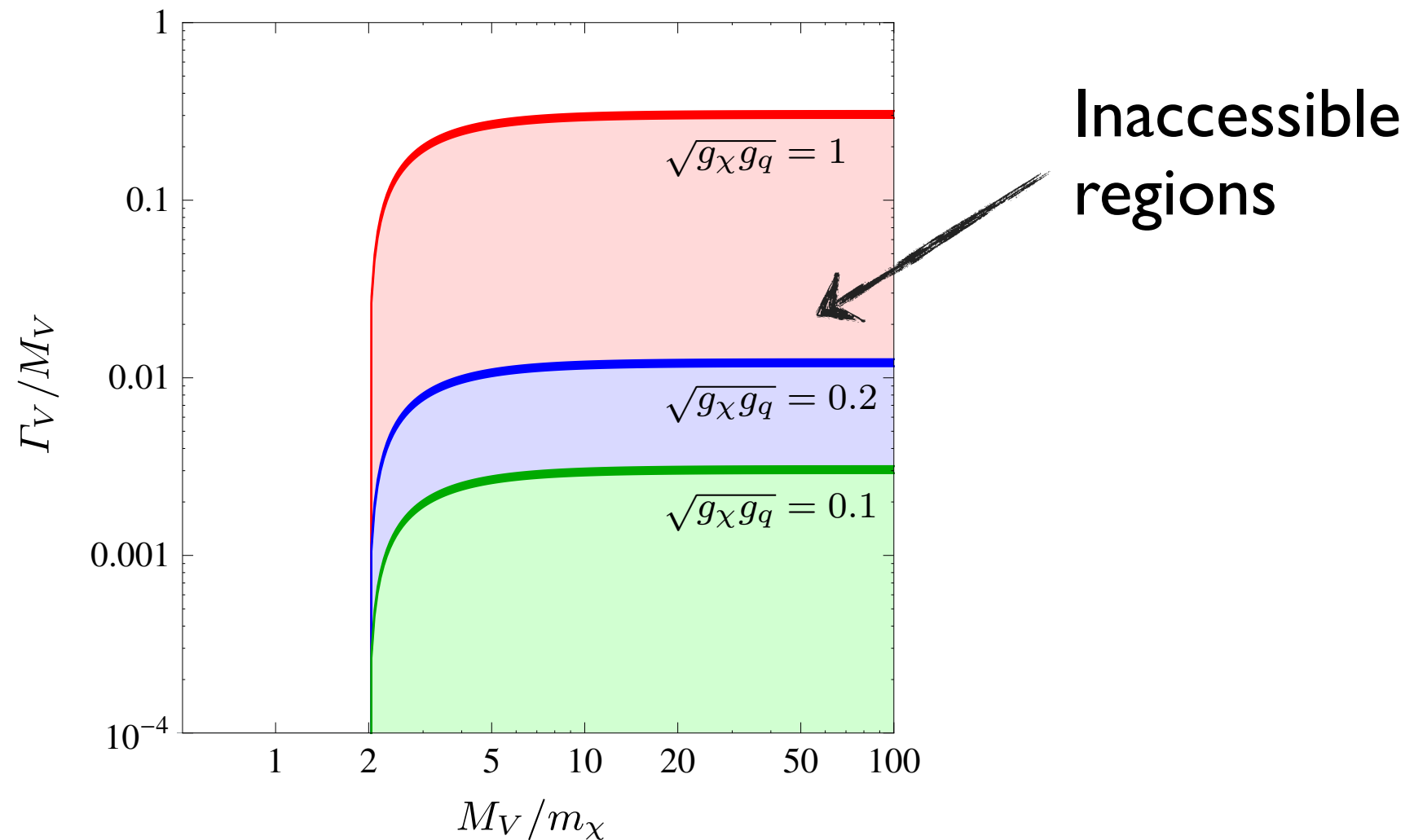
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( $\Rightarrow g_q, g_\chi$ )  
 EFT:  $m_\chi, (g_q g_\chi / M_V^2)$



# A "direct-detection-phobic" model



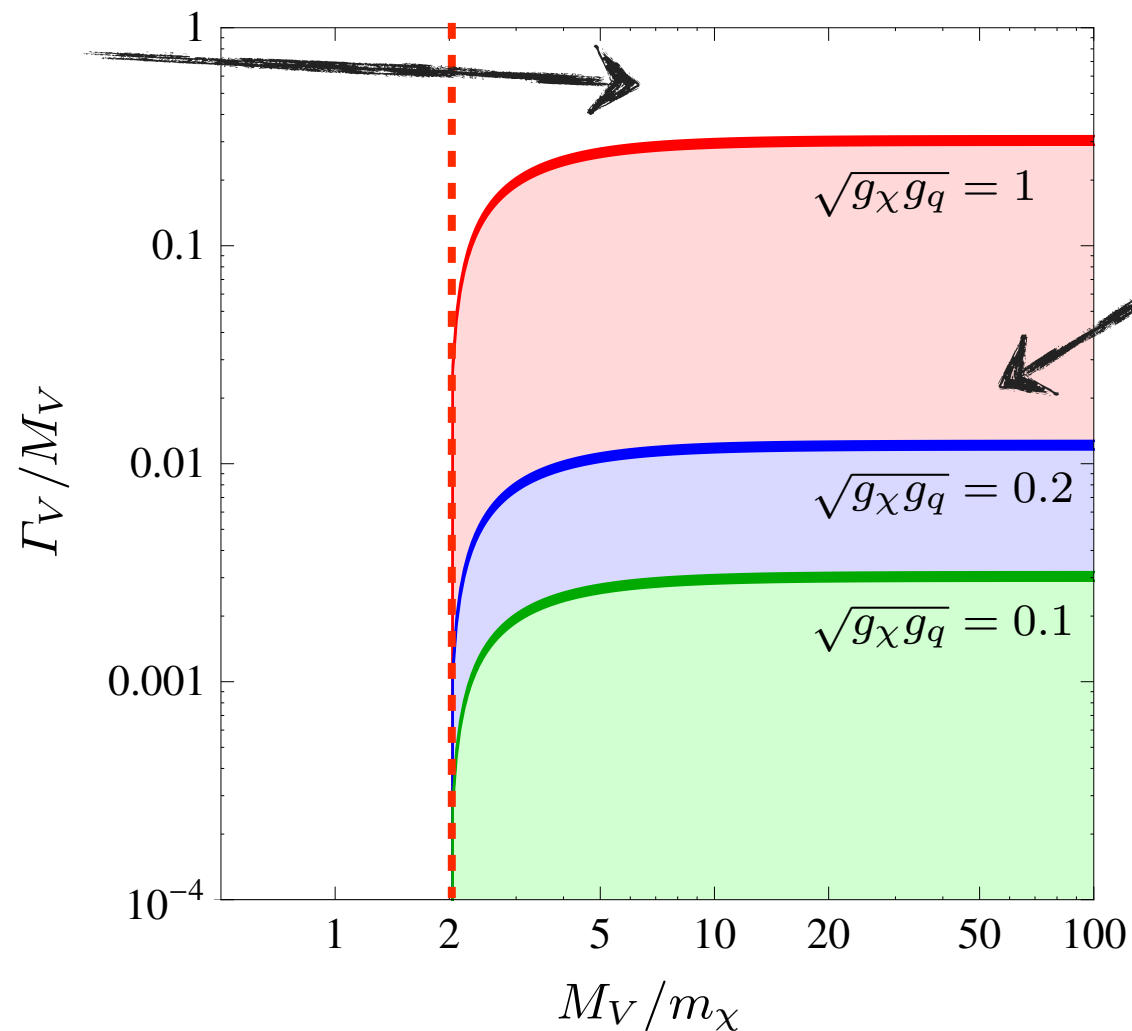
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# A "direct-detection-phobic" model

Two-fold ambiguity  
regarding  $g_q/g_\chi$

→ Choose  
smaller  $g_q/g_\chi$   
unless  $g_\chi > 4\pi$

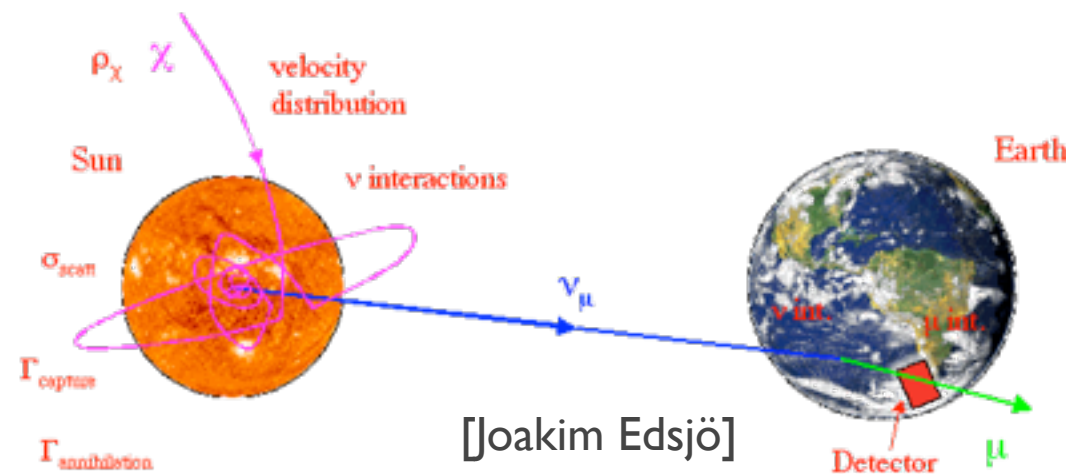


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Note: not all values allowed!

# IceCube limits from Dark Matter annihilation in the Sun

# Indirect DM detection: annihilation in the Sun

- Sun: Giant DM trap via WIMP-nucleon scattering ("direct detection")
- Sensitive to spin-dependent scattering



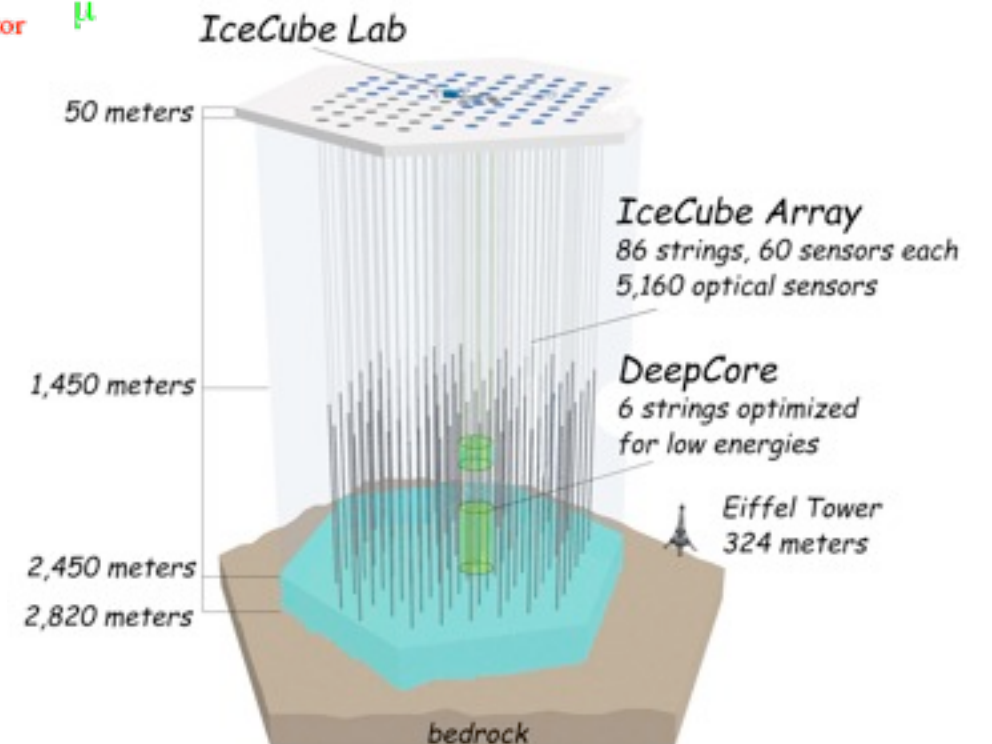
Detect neutrinos from DM annihilation with IceCube

$$\dot{N} = C_\odot - C_A N^2$$

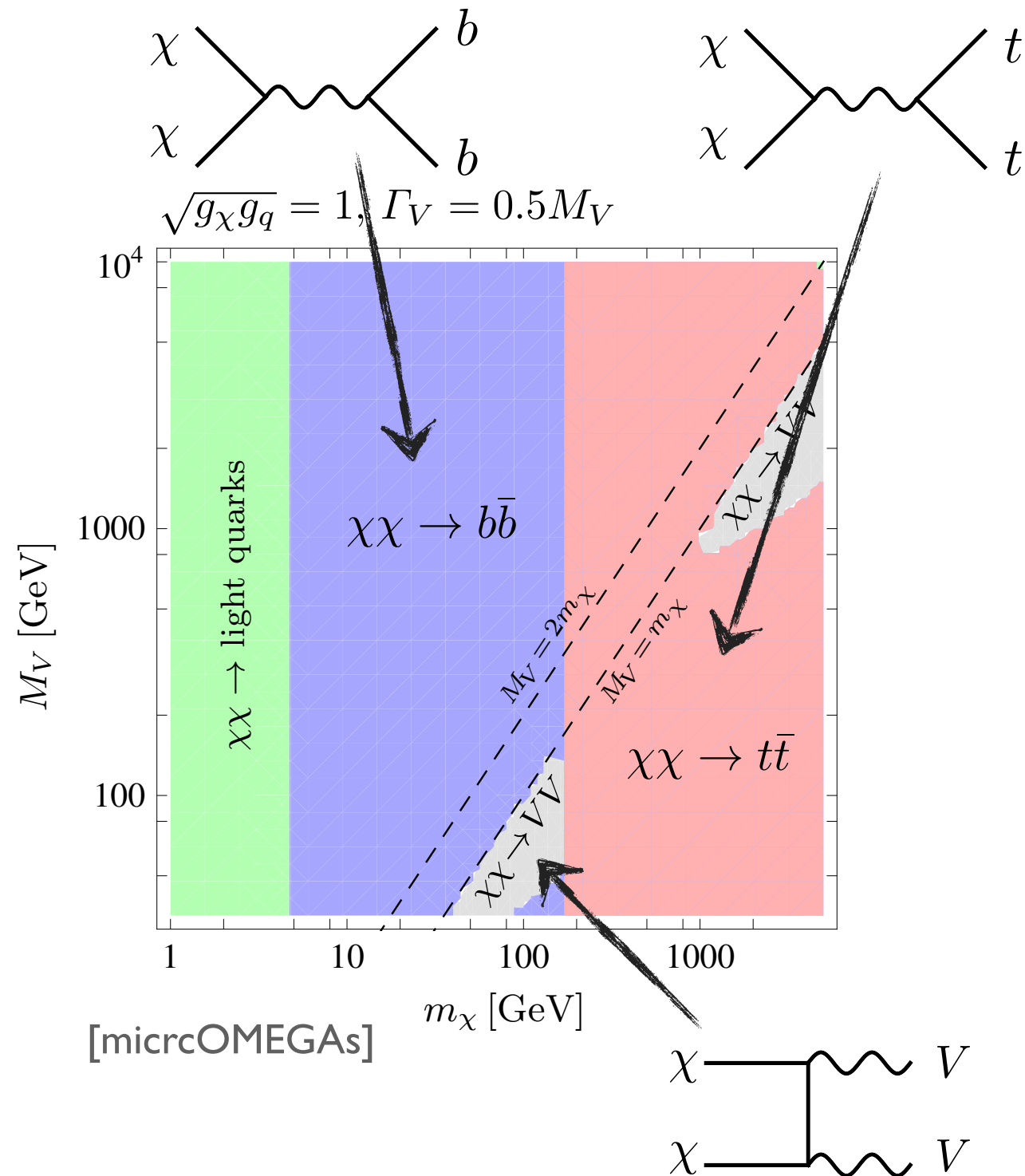
Capture rate:  
WIMP-nucleon  
scattering  $\sigma_{\text{SD}}$   
in the Sun

Annihilation rate:  
 $C_A N^2 = 2\Gamma_A$

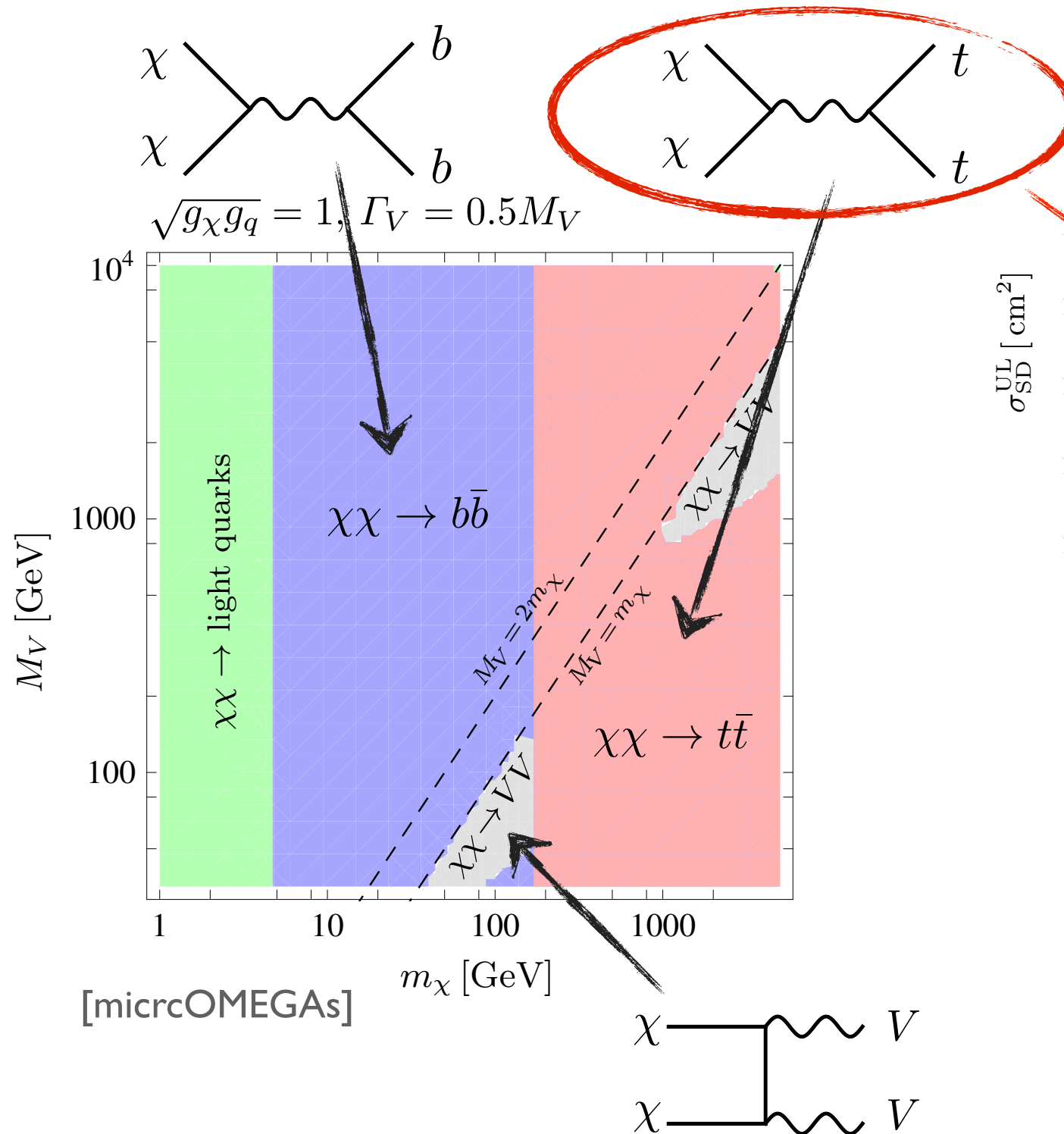
Equilibrium:  $C_\odot = C_A N^2 \propto$  neutrino flux



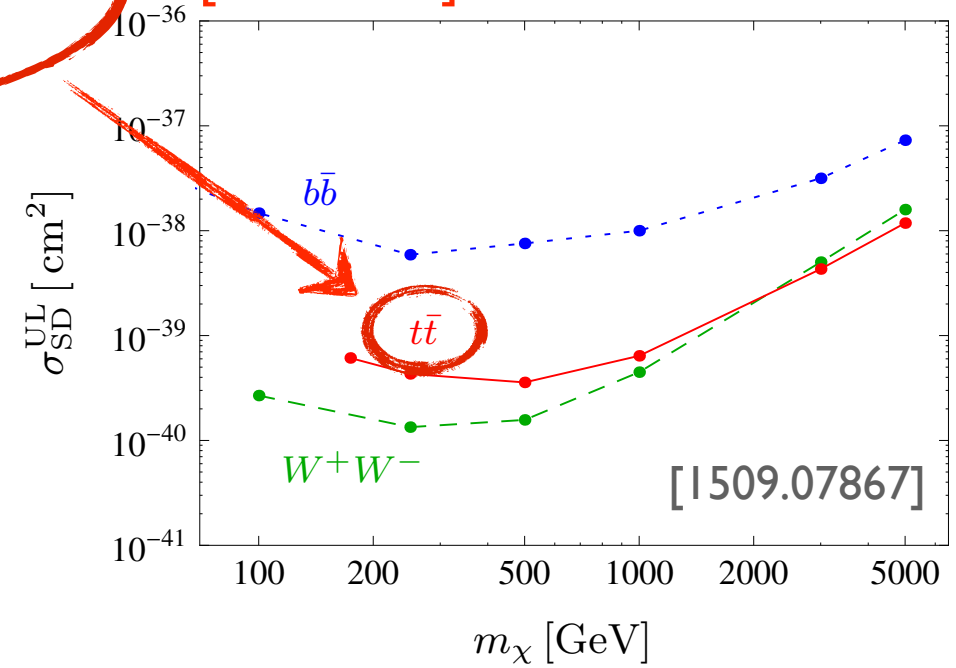
# Dominant annihilation channels



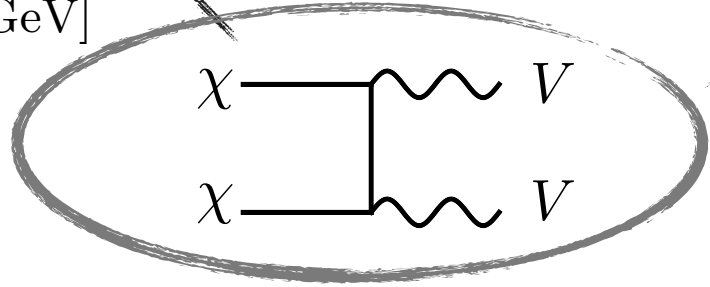
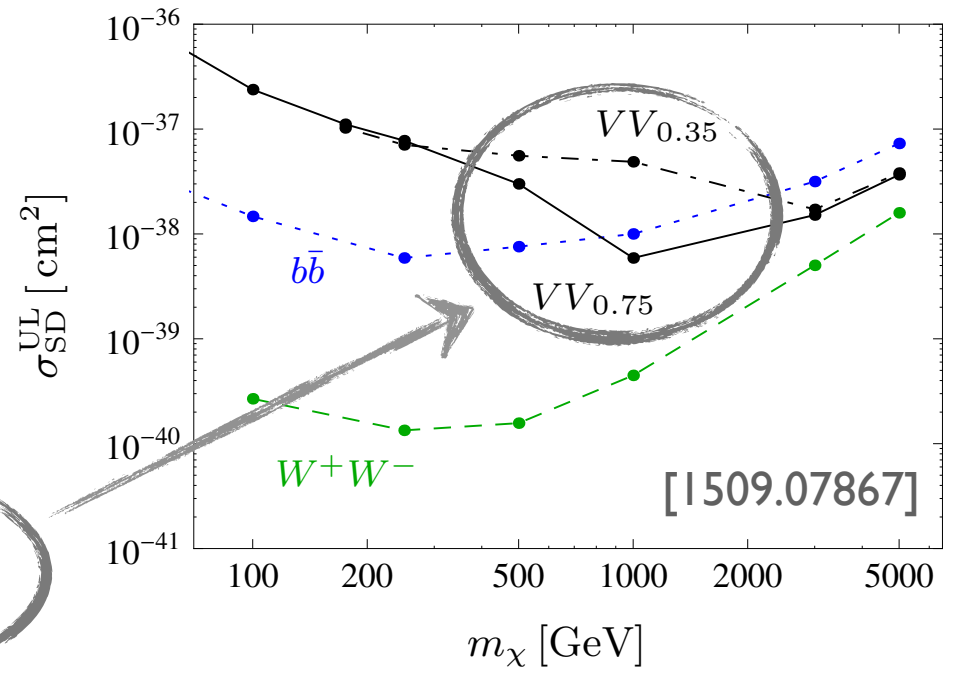
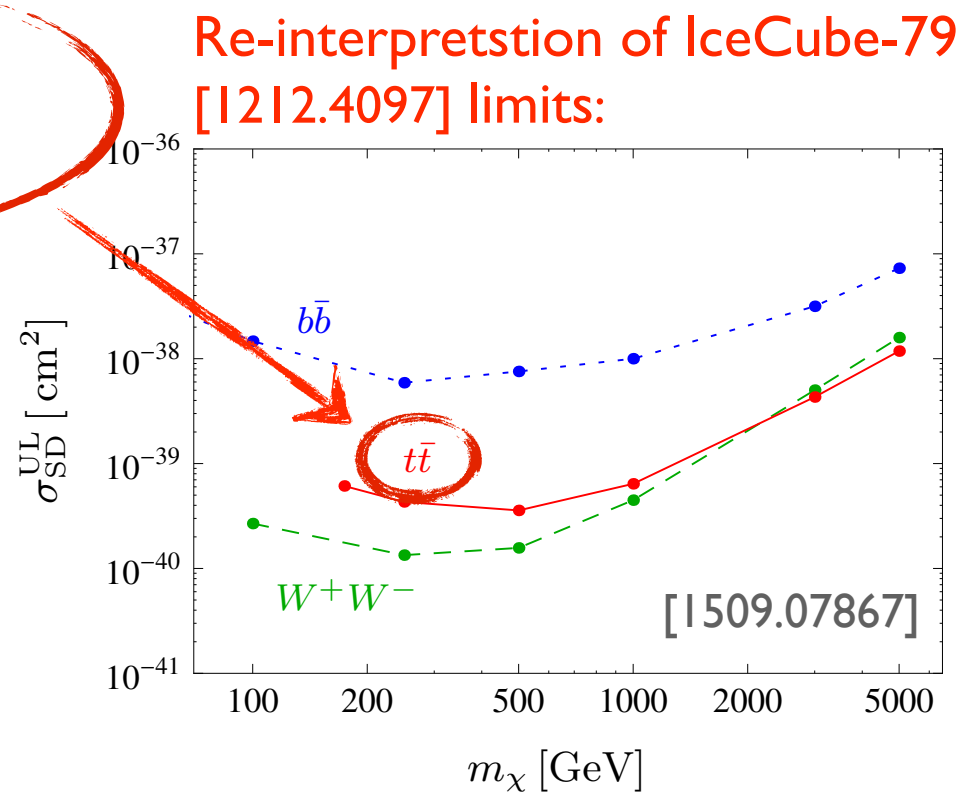
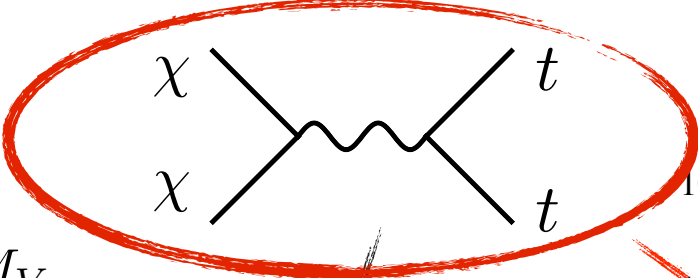
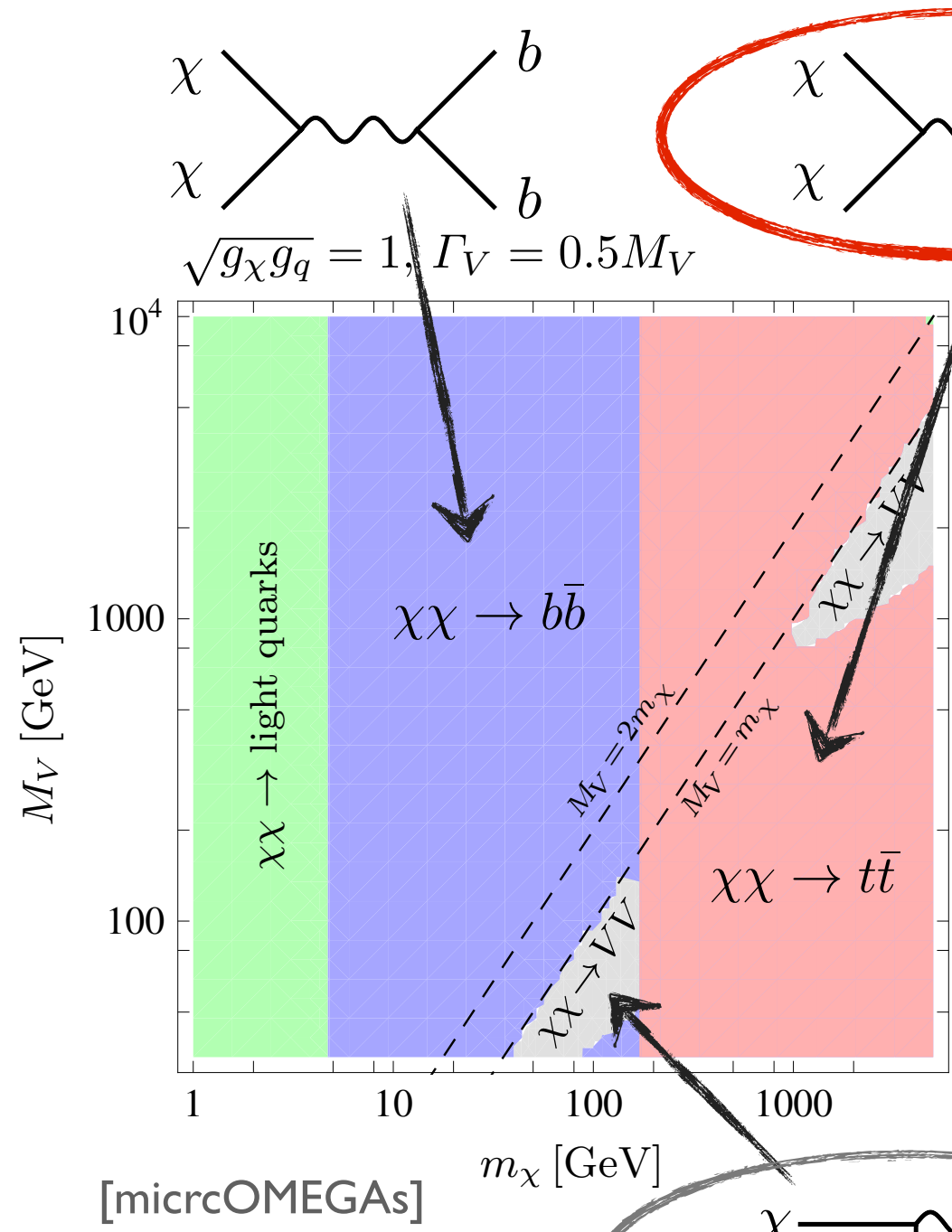
# Dominant annihilation channels



Re-interpretation of IceCube-79 [1212.4097] limits:



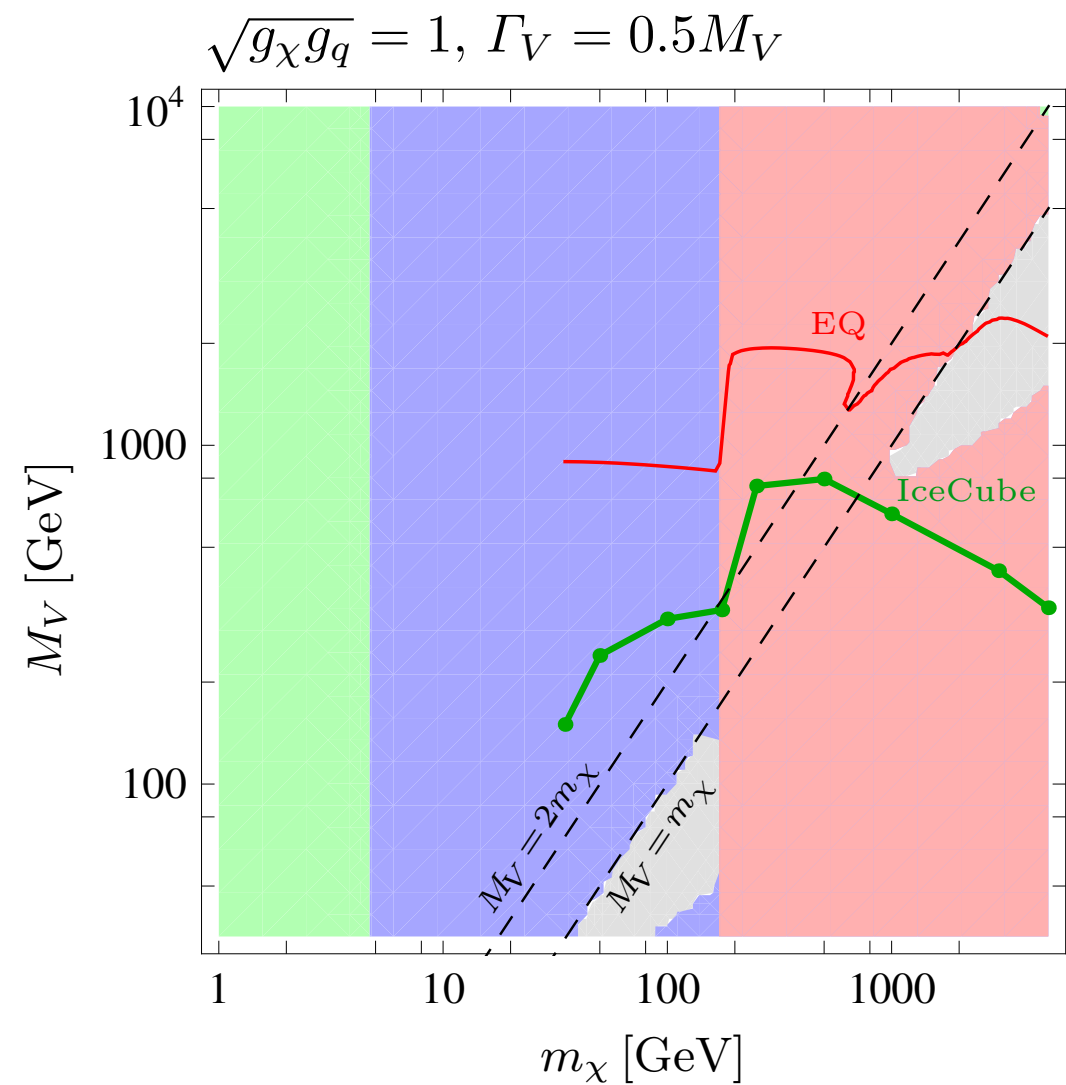
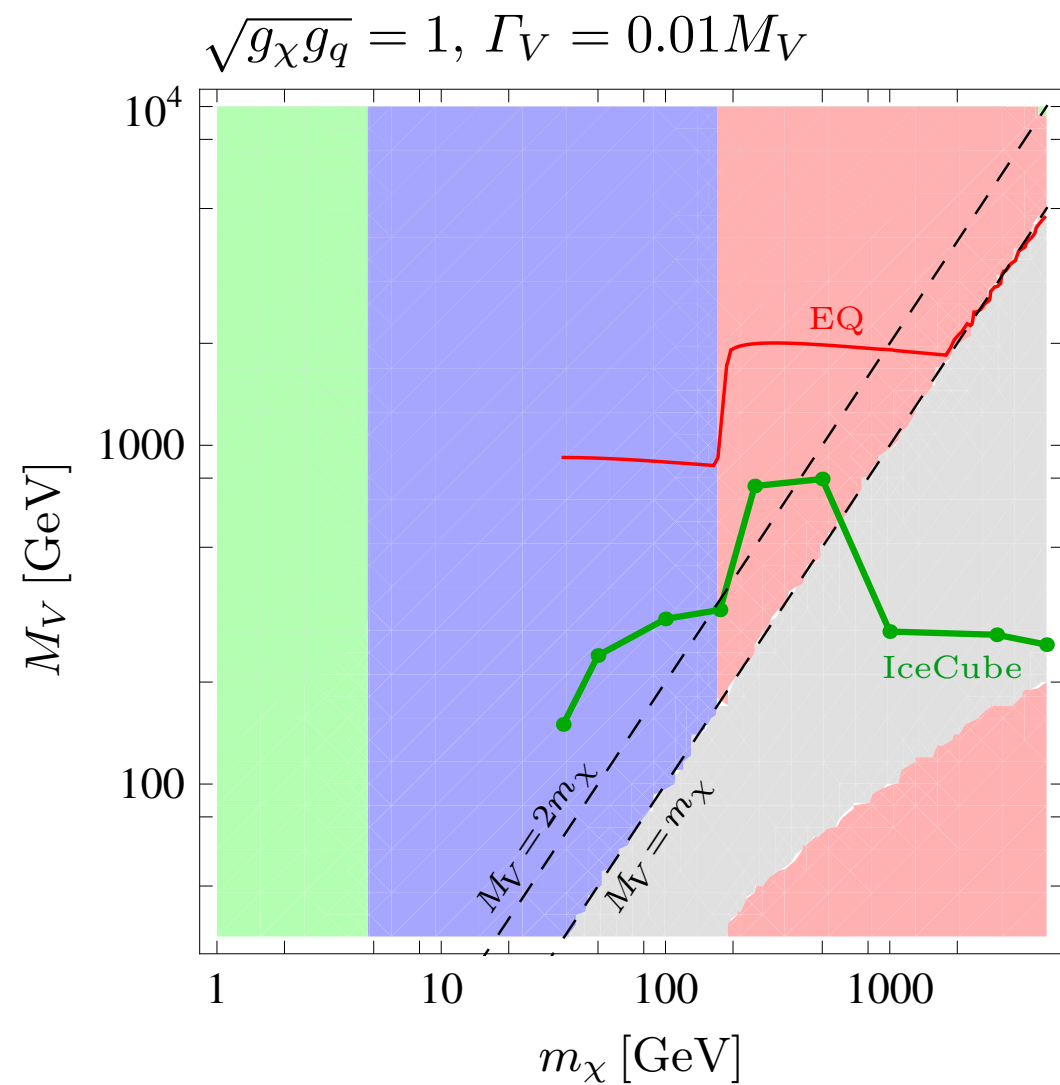
# Dominant annihilation channels





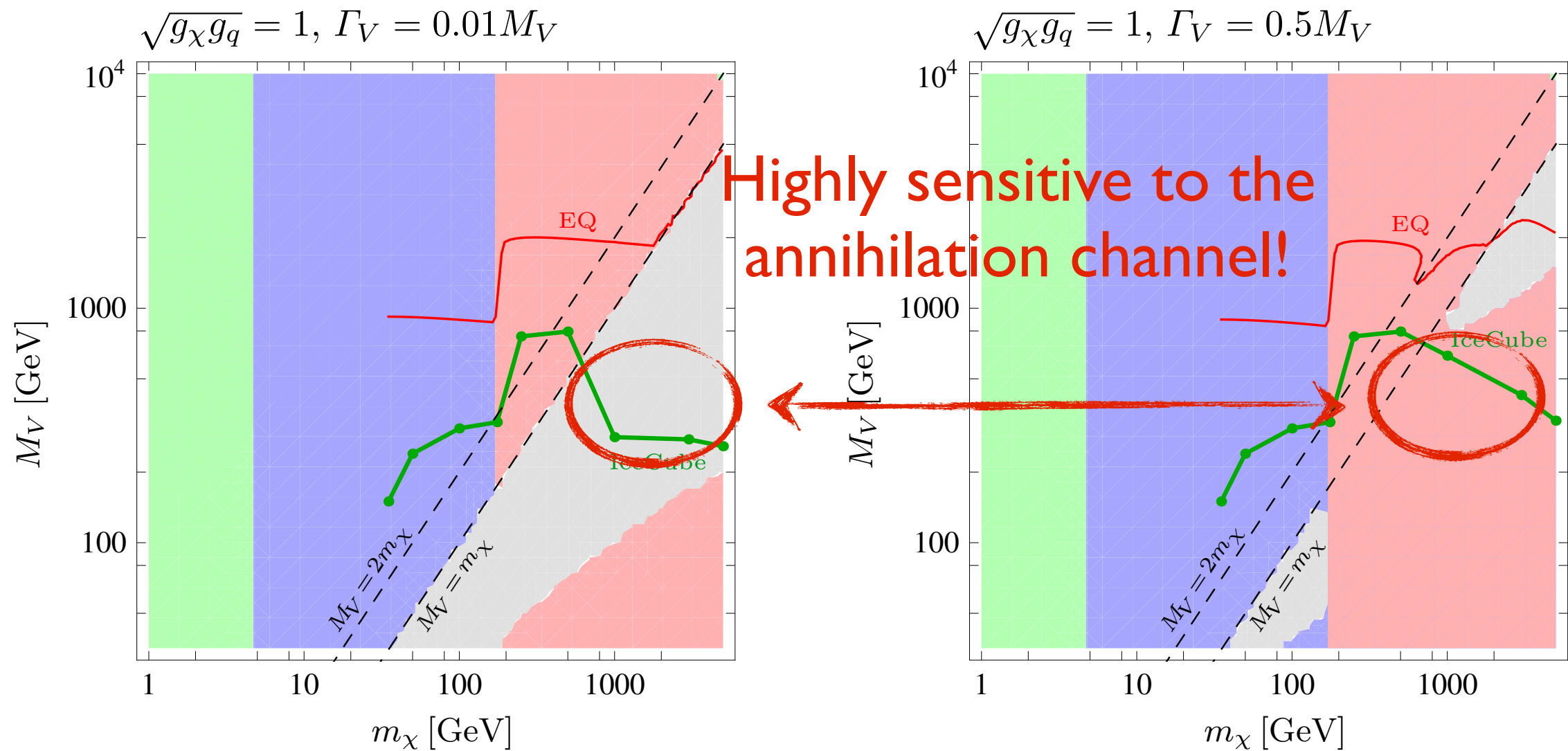
# Resulting limits from IceCube

[1509.07867]



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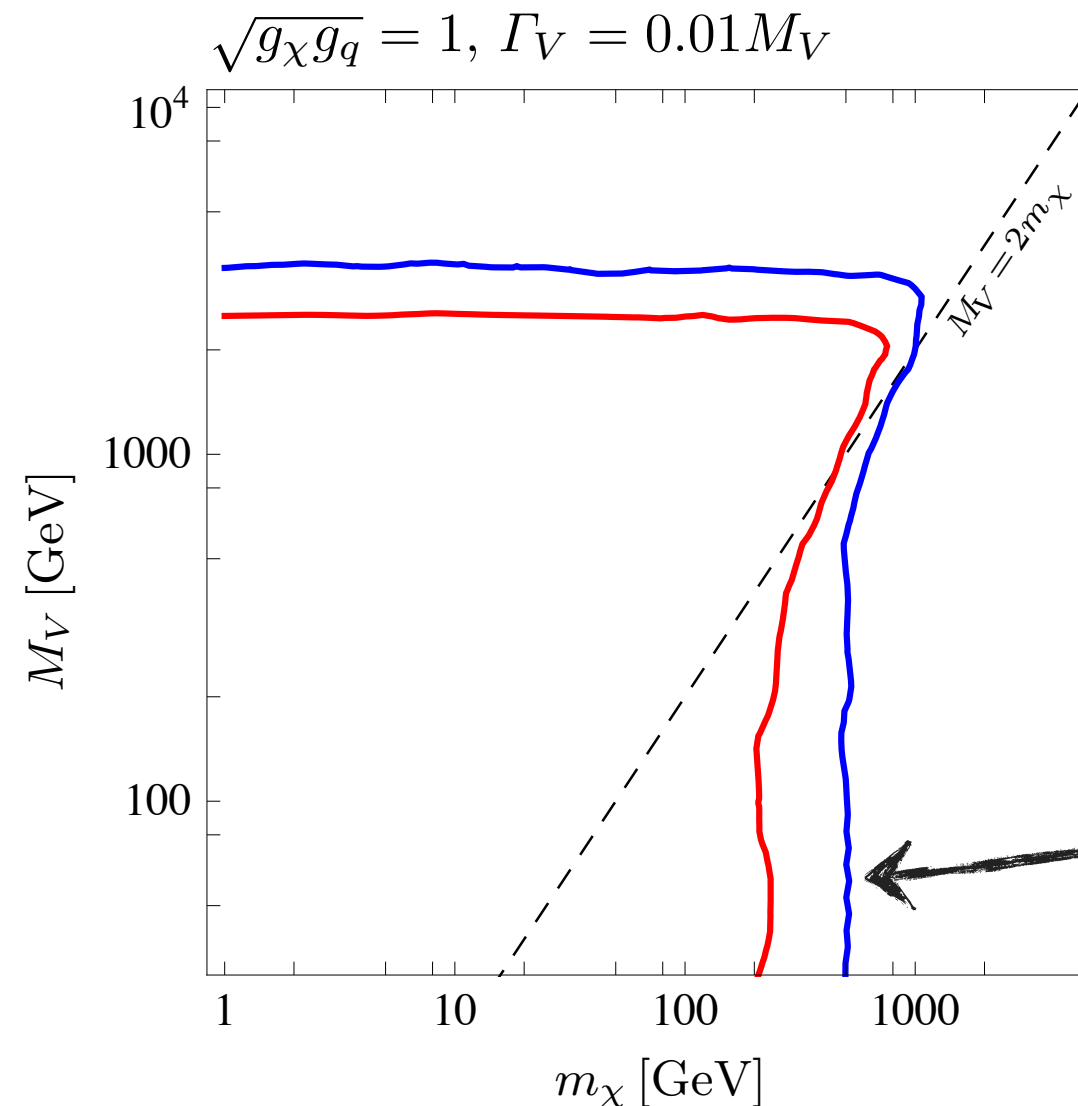
[1509.07867]



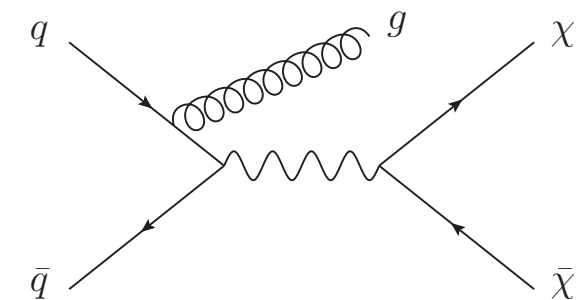
# Limits from the LHC

# Results from mono-jet searches at 8 TeV LHC

- Re-interpret LHC Run I mono-jet + MET searches  
[[ATLAS:1502.01518](#), [CMS:1408.3583](#)]
- Simulation: FeyRules/MadGraph/Pythia/Delphes

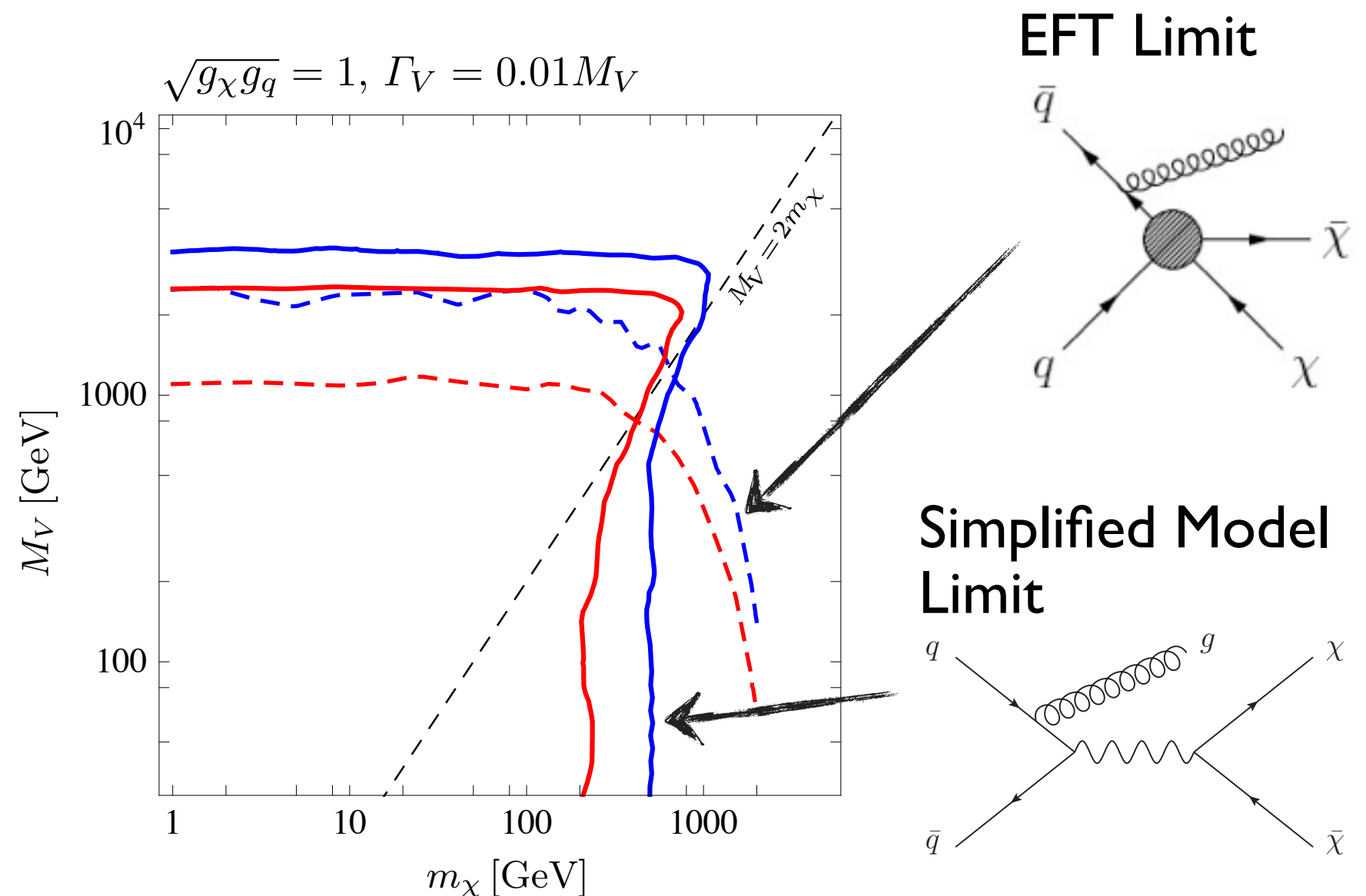


Simplified Model  
Limit



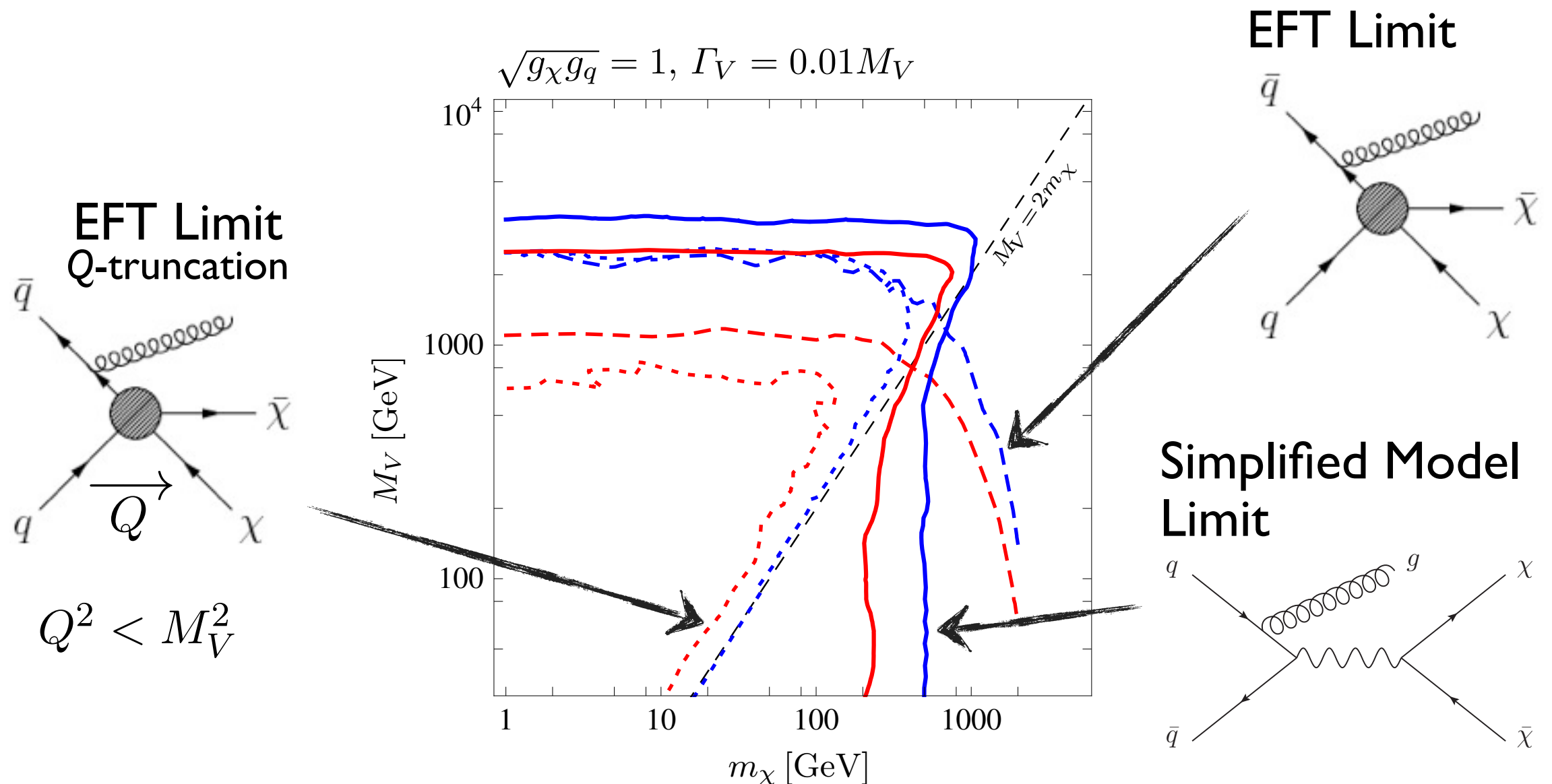
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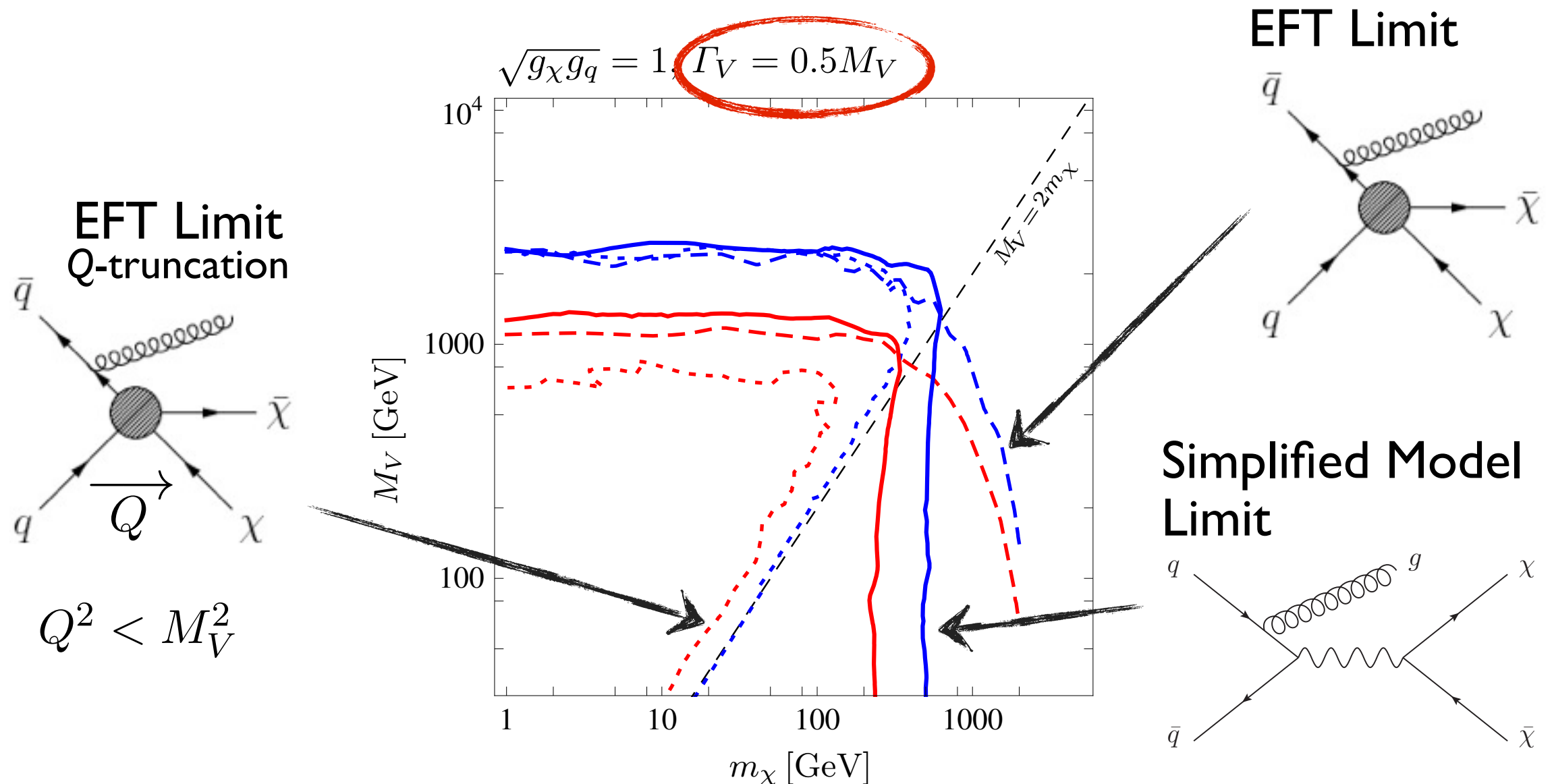
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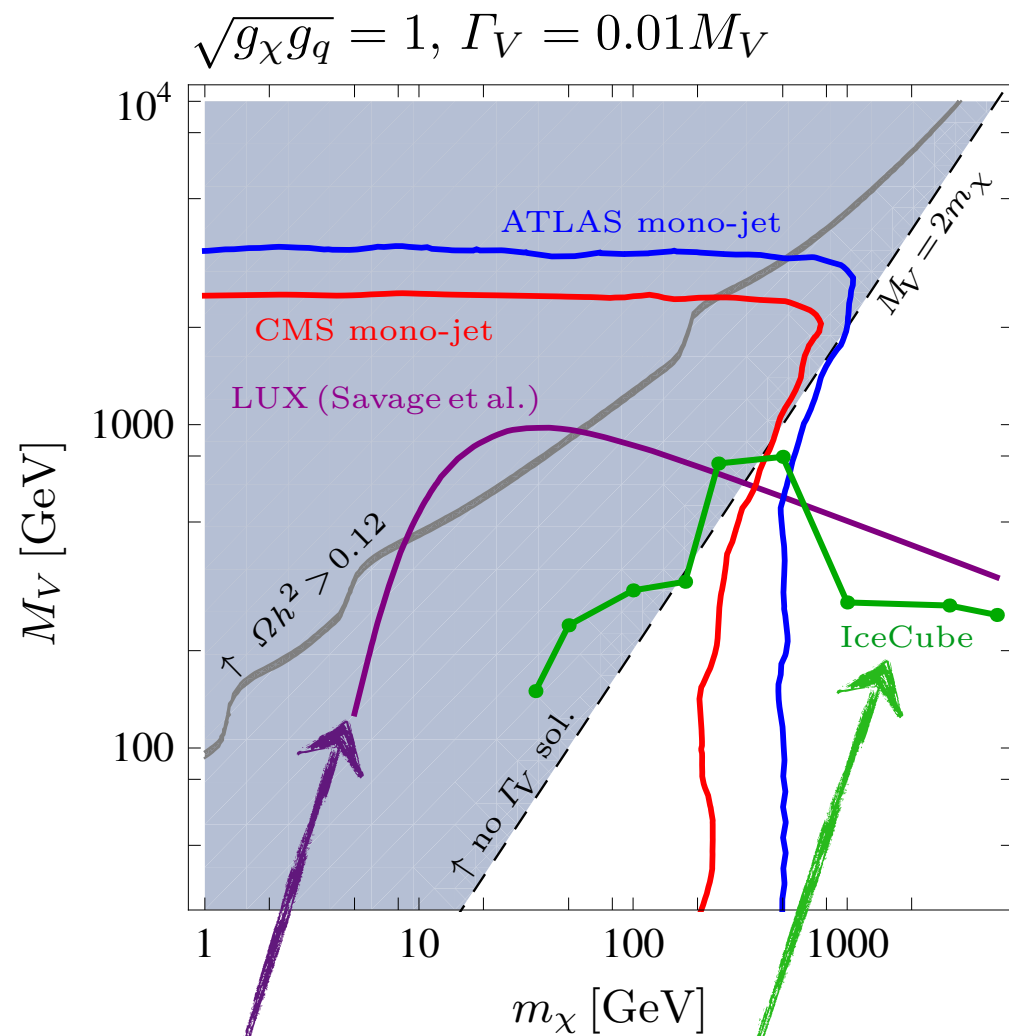




# Complementary constraints: Summary Plots

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[1509.07867]

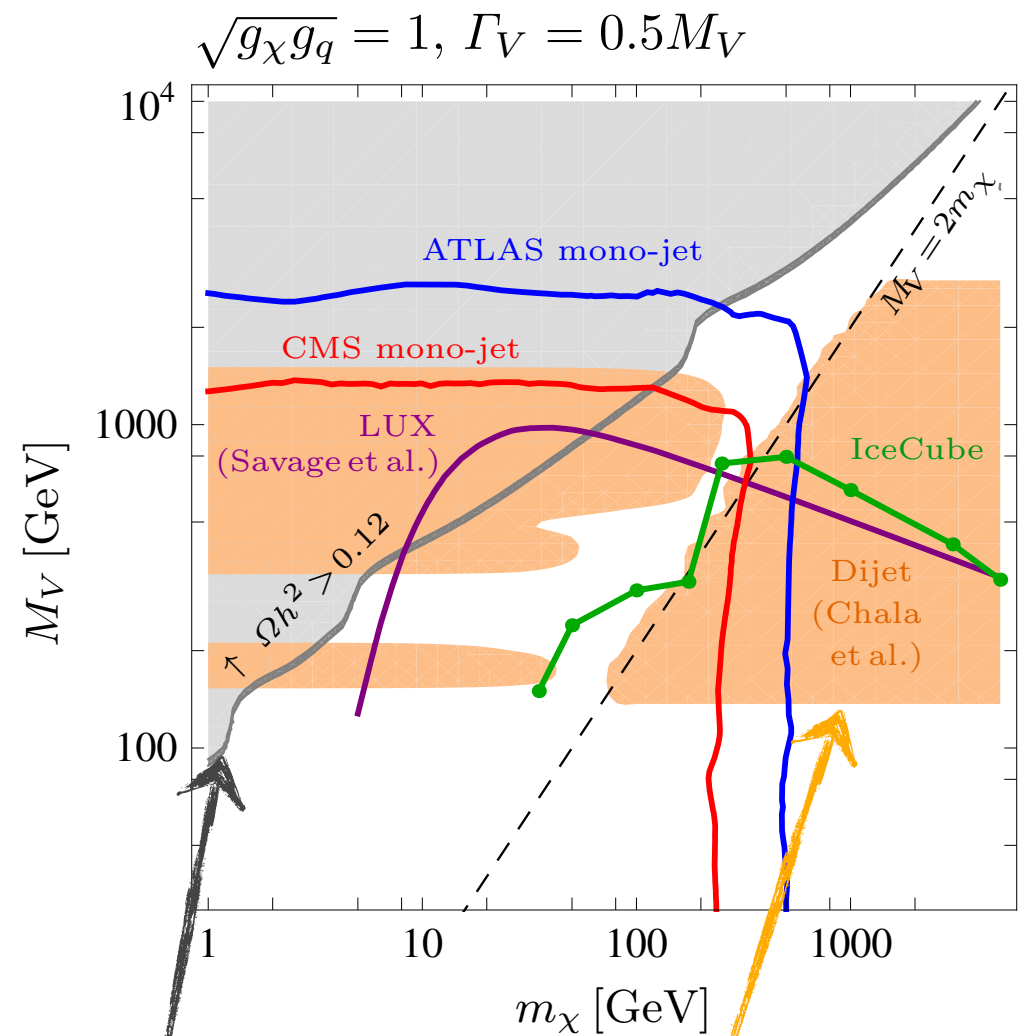
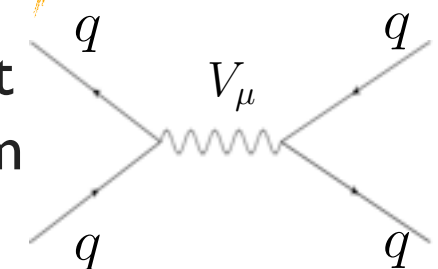


LUX limits  
from  
[1502.02667]

IceCube-79  
limits

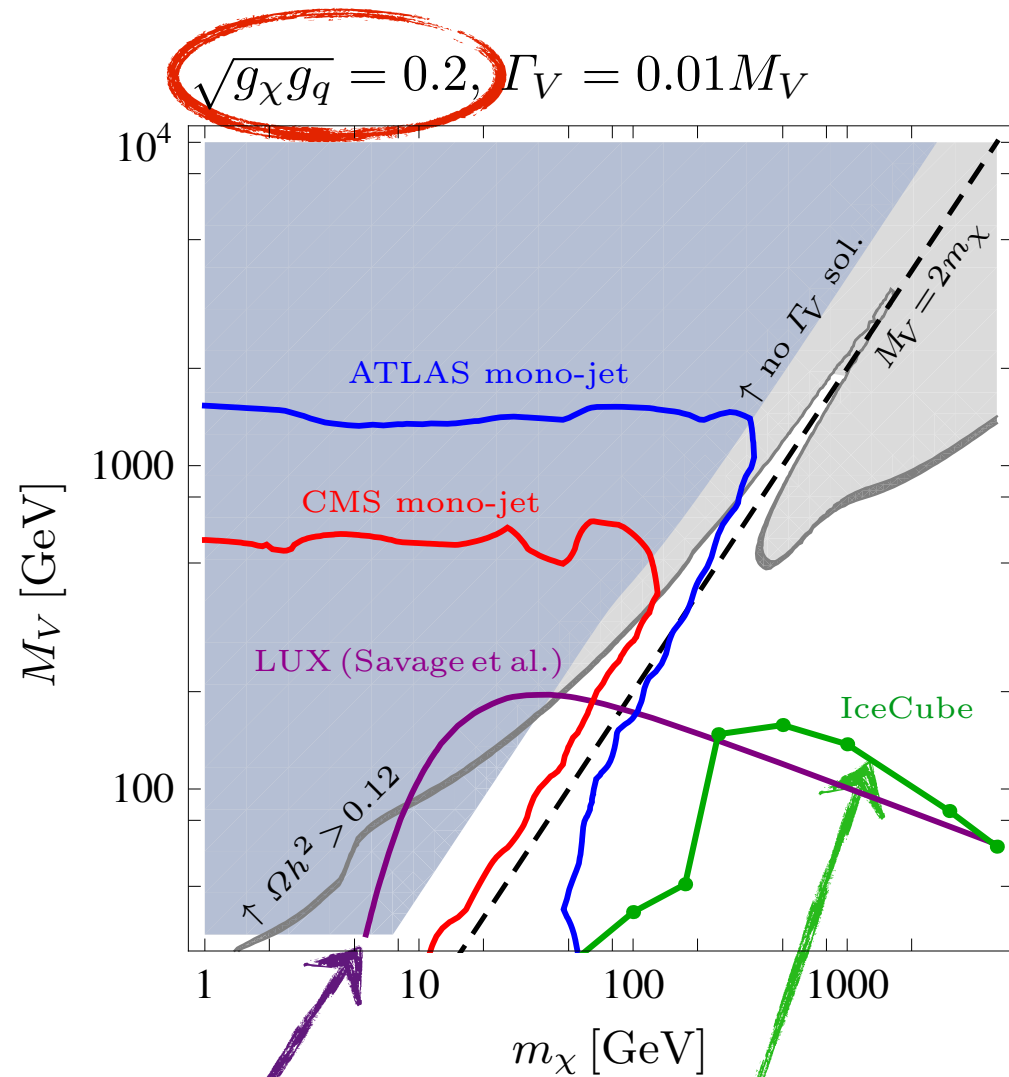
Relic density  
[micrOMEGAs]  
(simpl. model)

LHC Dijet  
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[1503.05916]



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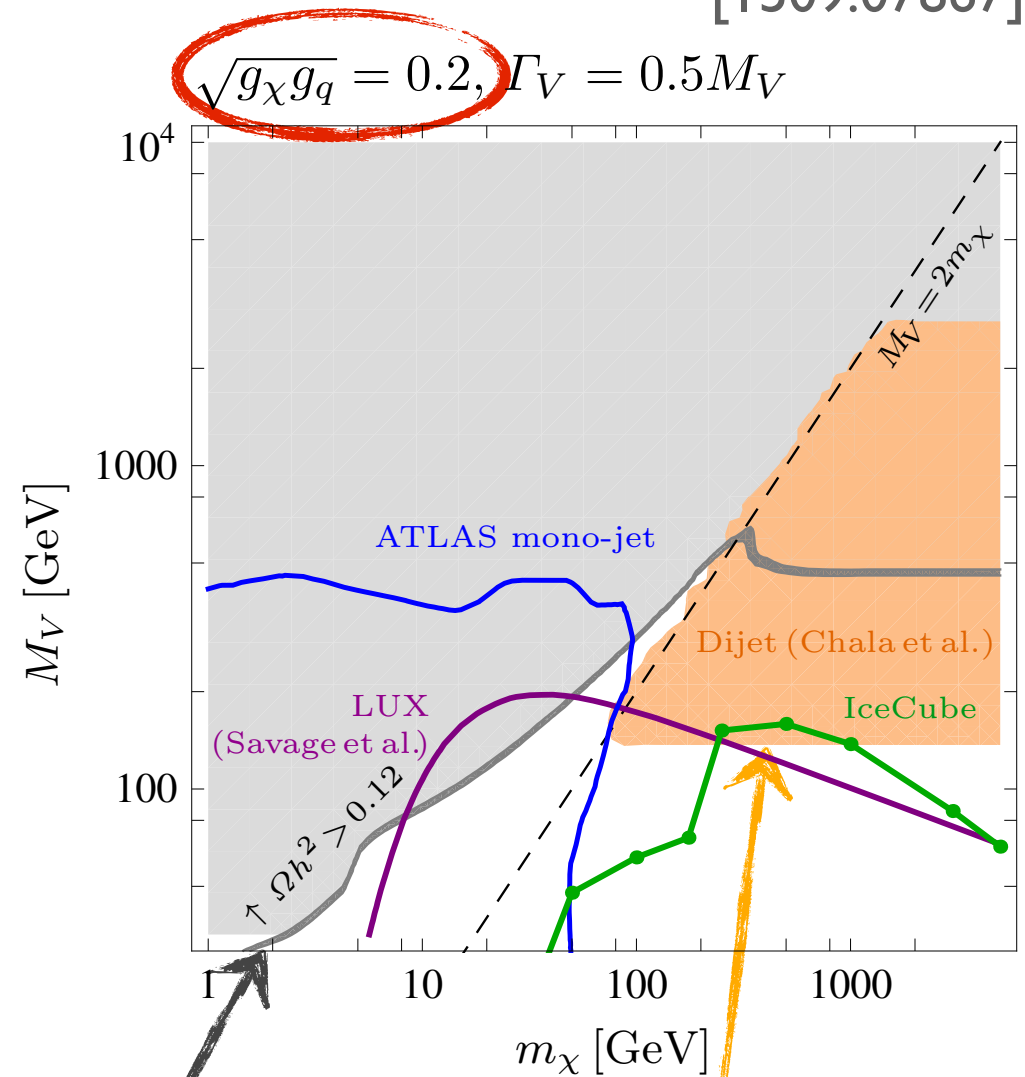
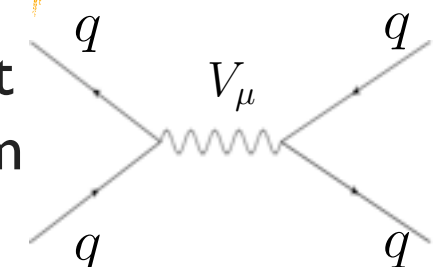


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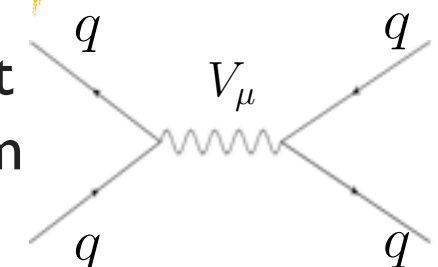
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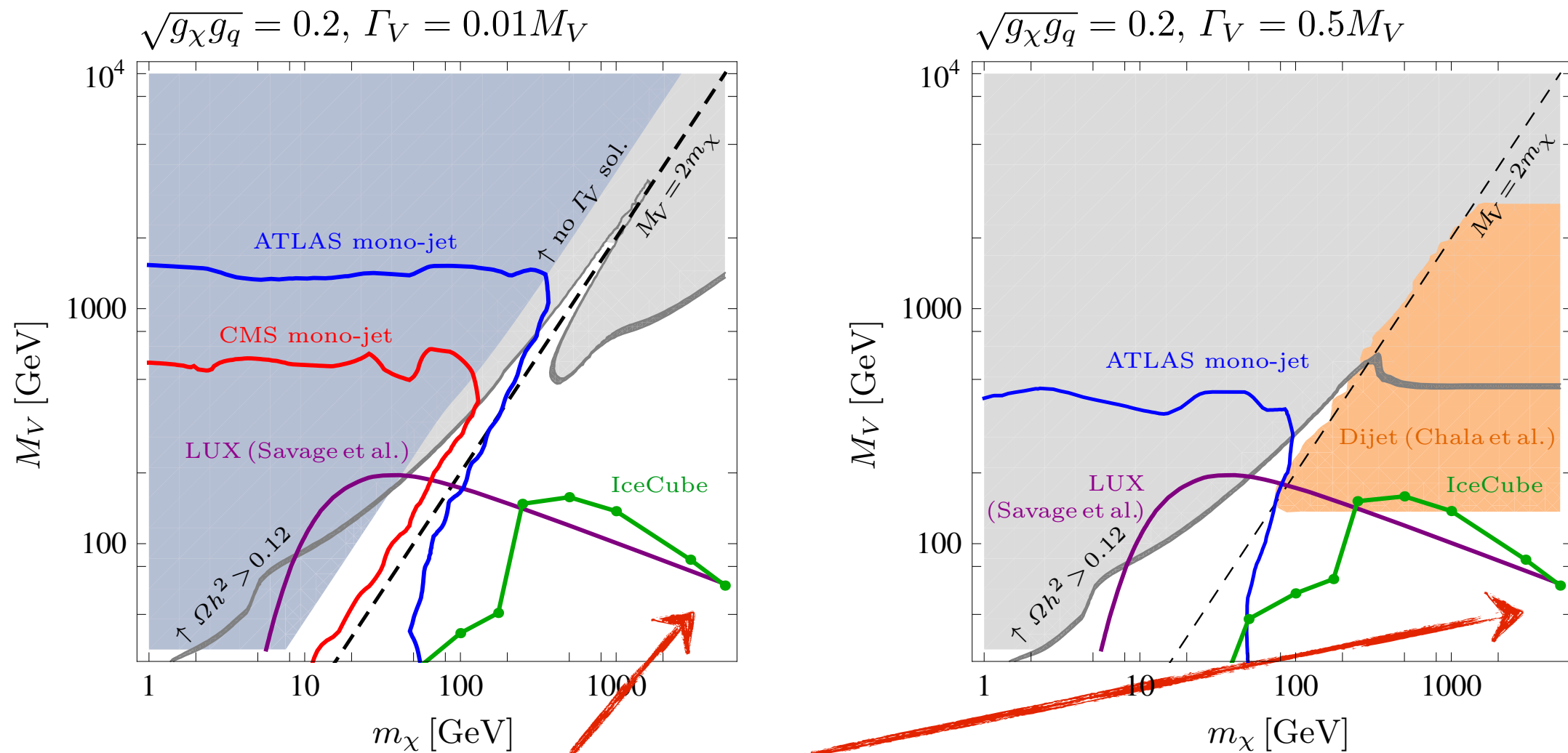
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# Complementary constraints: Summary

[1509.07867]



Note that  $M_V \ll m_\chi$  strongly constraint from unitarity [see e.g. 1510.02110]  
 → A more complete model is needed (add 'dark' Higgs sector)

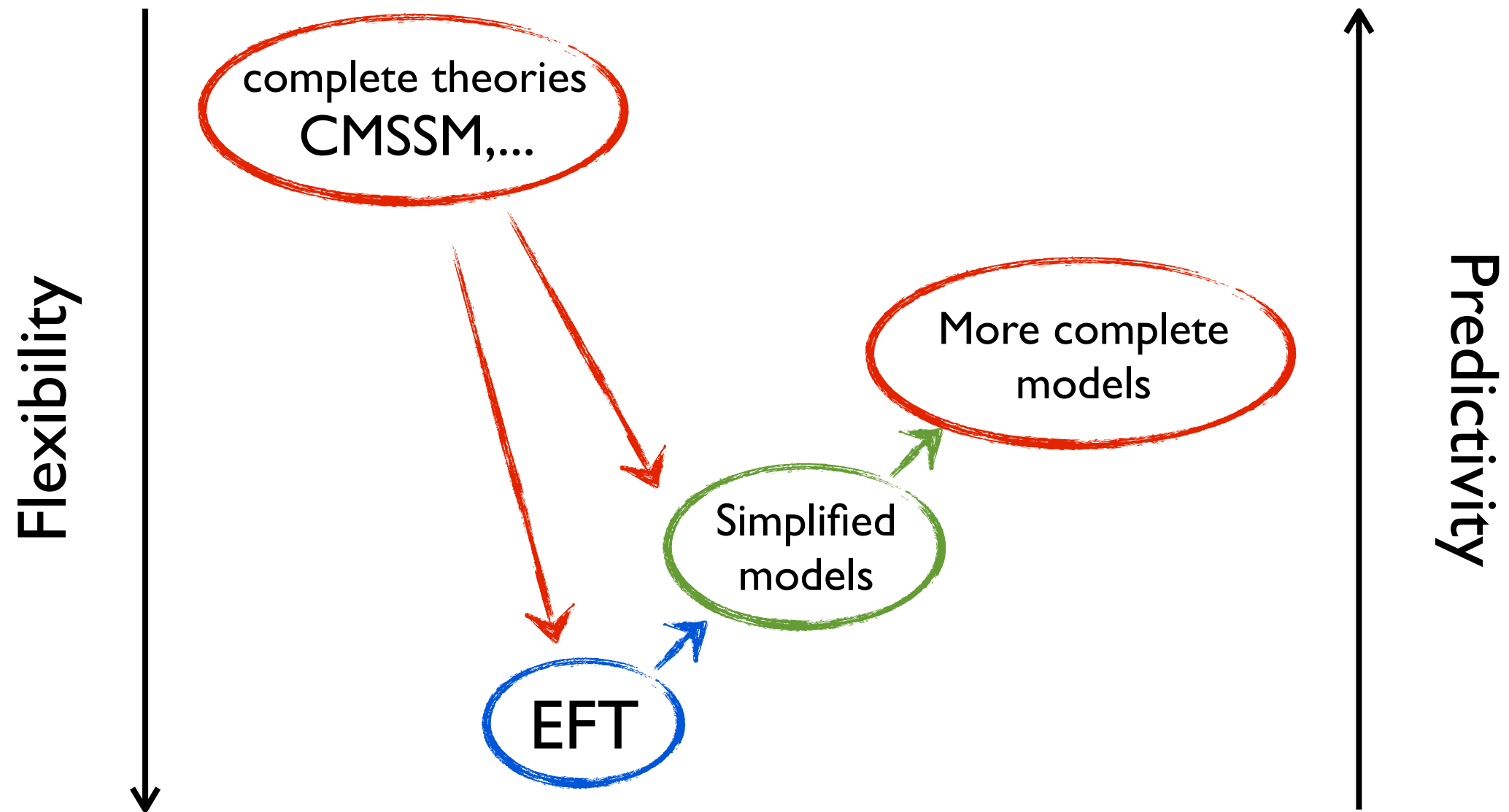
# Summary

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- Considered "direct detection-phobic" model: Vector mediator with axial couplings
  - Striking complementarity between various constraints
  - LHC: EFT not reliable,  $Q$ -truncation conservative estimate
  - ATLAS mono-jet strongest constraint on thermal relic strip sensitive up to  $M_V \simeq 3 \text{ TeV}$
  - IceCube important for annihilation into  $tt$ :  
Strongest limits for  $m_\chi \approx 200 - 500 \text{ GeV}$
-

# Summary

## Simplified and more complete models



**Thank you for your attention!**