

PERSPECTIVES FOR DETECTING LEPTON FLAVOUR VIOLATION IN LEFT-RIGHT SYMMETRIC MODELS

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Based on arXiv:1611.07025

in collaboration with Cesar Bonilla, Manuel E. Krauss and Werner Porod

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Theoretical Physics

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WHY LEFT-RIGHT SYMMETRIC MODELS?

MOTIVATION

① Automatic neutrino masses

- Combination both type-I and II seesaw

[R. Mohapatra & G. Senjanovic (1979), ...]

② GUT origin of gauge symmetries

[H. Fritzsch & P. Minkowski (1975), J. Pati & A. Salam (1973), C. Aulakh, et. al. (hep-ph/9712551), ...]

- $\text{SO}(10) \xrightarrow{M_{\text{GUT}}} (\mathcal{G}_{\text{PS}} \longrightarrow) \mathcal{G}_{\text{LR}} \xrightarrow{v_R} \mathcal{G}_{\text{SM}}$

$$\mathcal{G}_{\text{PS}} = \text{SU}(4) \times \text{SU}(2)_L \times \text{SU}(2)_R$$

$$\mathcal{G}_{\text{LR}} = \text{SU}(3)_C \times \text{SU}(2)_L \times \text{SU}(2)_R \times \text{U}(1)_{B-L}$$

③ Strong-CP problem

[A. Maiezza & M. Nemevšek (arXiv:1407.3678), ...]

- Potentially solved by \mathcal{C} & \mathcal{P} restoration

So, WHAT'S THE PLAN?

① Preliminaries

② Parametrising the neutrino sector

- Casas-Ibarra parametrisation
- Our parametrisation

③ Lepton flavour violation results

- Current and future bounds
- (Effect of δ_{CP})

PRELIMINARIES

FIELD CONTENT, YUKAWA SECTOR AND ALL THAT JAZZ!

Quick-fire recap from last time:

$$\mathcal{L}_Y = \mathcal{L}_Y^\Phi + \mathcal{L}_Y^\Delta + \text{h.c.}$$

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$$\mathcal{L}_Y = \mathcal{L}_Y^\Phi + \mathcal{L}_Y^\Delta + \text{h.c.}$$

$$-\mathcal{L}_Y^\Phi = \overline{Q_L} \left(Y_{Q_1} \Phi + Y_{Q_2} \tilde{\Phi} \right) Q_R \\ + \overline{L_L} \left(Y_{L_1} \Phi + Y_{L_2} \tilde{\Phi} \right) L_R$$

Field	Gen.	\mathcal{G}_{LR}
Q_L	3	$(\mathbf{3}, \mathbf{2}, \mathbf{1}, \frac{1}{3})$
Q_R	3	$(\mathbf{3}, \mathbf{1}, \mathbf{2}, \frac{1}{3})$
L_L	3	$(\mathbf{1}, \mathbf{2}, \mathbf{1}, -1)$
L_R	3	$(\mathbf{1}, \mathbf{1}, \mathbf{2}, -1)$
Φ	1	$(\mathbf{1}, \mathbf{2}, \mathbf{2}, 0)$

$$\tilde{\Phi} \equiv -\sigma_2 \Phi^* \sigma_2$$

$$\overline{\Psi^C} = \Psi^T C, \quad C = i\gamma_2\gamma_0$$

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$$\begin{aligned}\mathcal{L}_Y &= \mathcal{L}_Y^\Phi + \mathcal{L}_Y^\Delta + \text{h.c.} \\ -\mathcal{L}_Y^\Phi &= \overline{Q_L} \left(Y_{Q_1} \Phi + Y_{Q_2} \tilde{\Phi} \right) Q_R \\ &\quad + \overline{L_L} \left(Y_{L_1} \Phi + Y_{L_2} \tilde{\Phi} \right) L_R \\ -\mathcal{L}_Y^\Delta &= \overline{L_L^C} Y_{\Delta_L} (i\sigma_2) \Delta_L L_L \\ &\quad + \overline{L_R^C} Y_{\Delta_R} (i\sigma_2) \Delta_R L_R\end{aligned}$$

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↪ SARAH model file publicly available

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DETERMINING THE NEUTRINO SECTOR

NEUTRINO MASSES

Neutrino mass matrix from \mathcal{L}_Y

$$\mathcal{M}_\nu = \begin{pmatrix} M_L^* & M_D \\ M_D^T & M_R \end{pmatrix}$$

$$\begin{aligned} M_{L/R} &= \sqrt{2} Y_{\Delta_{L/R}} v_{L/R} \\ M_D &= \frac{1}{\sqrt{2}} (Y_{L_1} v_1 + Y_{L_2} v_2) \end{aligned}$$

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IN THE SEESAW APPROXIMATION

$$m_\nu^{\text{light}} = U_{\text{PMNS}}^* \text{diag}(m_{\nu_1}, m_{\nu_2}, m_{\nu_3}) U_{\text{PMNS}}^\dagger \stackrel{!}{=} (M_L^* - M_D M_R^{-1} M_D^T)$$

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How to determine model parameters?

- Casas-Ibarra parametrisation
- Brute force fits
- Our parametrisation

CASAS-IBARRA PARAMETRISATION

Simple procedure to find M_D :

$$M_D = f(m_\nu^{\text{light}}, M_L, M_R, \textcolor{red}{R})$$

↪ Require an additional matrix $\textcolor{red}{R}$

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$$M_D = -iV^* D_{X_\nu}^{1/2} R D_{M_R}^{1/2} O^\dagger$$

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IMPLICATIONS:

- R is an unknown matrix $\in O(3; \mathbb{C})$
- Unrelated to model parameters
- Potentially violates \mathcal{C} & \mathcal{P}
- Results in an infinite number of degenerate solutions

DISCRETE SYMMETRIES

Parity \mathcal{P} :

$$\left. \begin{array}{l} L_L \leftrightarrow L_R \\ \Delta_L \leftrightarrow \Delta_R \\ \Phi \leftrightarrow \Phi^\dagger \end{array} \right\} \xrightarrow{\mathcal{L} \text{ inv.}} \left\{ \begin{array}{l} Y_{Q_i} = Y_{Q_i}^\dagger \\ Y_{L_i} = Y_{L_i}^\dagger \\ Y_{\Delta_L} = Y_{\Delta_R} \end{array} \right.$$

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CONSEQUENCE:

$$m_\nu^{\text{light}} \stackrel{\mathcal{P}}{=} \left(\frac{v_L}{v_R} M_R^* - M_D M_R^{-1} M_D^* \right)$$

$$m_\nu^{\text{light}} \stackrel{\mathcal{C}}{=} \left(\frac{v_L}{v_R} M_R - M_D M_R^{-1} M_D \right)$$

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DIRECT ALGEBRAIC SOLUTIONS

$$Y_\Delta \equiv Y_{\Delta_L}^{(*)} = Y_{\Delta_R} = f(M_D, v_L, v_R)$$

[E. Akhmedov and M. Frigerio ([hep-ph/0509299](#))]

OUR PARAMETRISATION

\mathcal{C} -conserving case re-written as:

$$\begin{aligned} B &= \alpha A - A^{-1} \\ &= V^* (\alpha D_A - D_A^{-1}) V^\dagger \end{aligned}$$

$$\begin{aligned} B &\equiv M_D^{-1/2} m_\nu^{\text{light}} M_D^{-1/2}, \quad \alpha \equiv \frac{v_L}{v_R} \\ A &\equiv M_D^{-1/2} M_R M_D^{-1/2} \end{aligned}$$

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Diagonalise and solve resulting diagonal polynomial matrix equations

$$D_A^{(i,i)} = \frac{D_B^{(i,i)} \pm \sqrt{\left(D_B^{(i,i)}\right)^2 + 4\alpha}}{2\alpha}$$

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RESULT

$$Y_\Delta^{(\pm\pm\pm)} = \frac{1}{2\sqrt{2}v_L} M_D^{1/2} V^* \text{diag} \left(D_B^{(i,i)} \pm \sqrt{\left(D_B^{(i,i)}\right)^2 + 4\alpha} \right) V^\dagger M_D^{1/2}$$

Advantages:

- ➊ Eightfold degenerate solutions
- ➋ Function only of model parameters

LEPTON FLAVOUR VIOLATION

LEPTON FLAVOUR OBSERVABLES

CALCULATED OBSERVABLES:

- $\ell_\alpha \rightarrow \ell_\beta \gamma$
- $\ell_\alpha \rightarrow 3\ell_\beta$
- $\ell_\alpha \rightarrow \ell_\beta \ell_\gamma \ell_\delta$
- $\mu - e$ conversion

Including all tree and one-loop contributions

LEPTON FLAVOUR OBSERVABLES

CALCULATED OBSERVABLES:

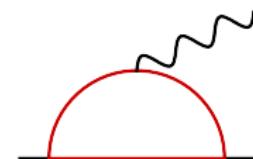
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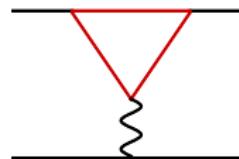
Example Topologies



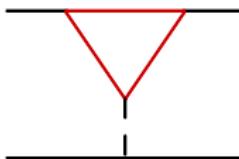
Tree-level scalar



Radiative



Vector penguin



Scalar penguin



Box

LEPTON FLAVOUR OBSERVABLES

KEY PARAMETERS (AS A RESULT OF THE PARAMETRISATION)

- v_L
- δ_{CP}
- M_D
 - (i) $x1 \text{ GeV}$
 - (ii) $xM_{\text{up-type}}$
 - (iii) $xV_{\text{CKM}}^\dagger M_{\text{up-type}} V_{\text{CKM}}^*$
- Y_Δ
 - (i) $Y_\Delta^{(+++)}$
 - (ii) $Y_\Delta^{(+--)}$

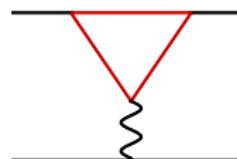
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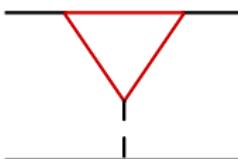
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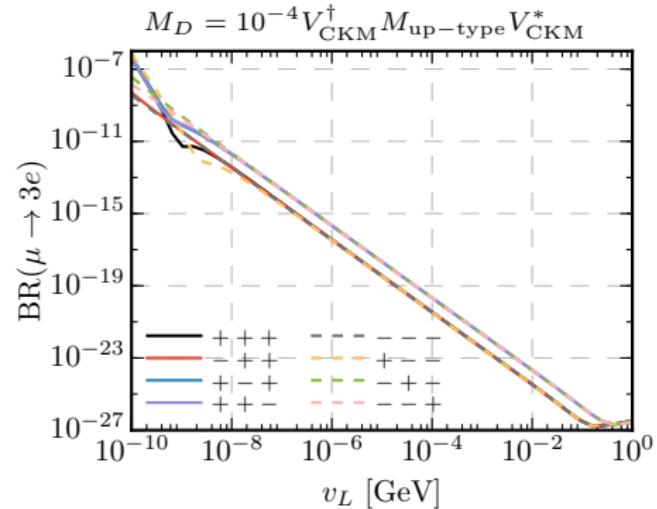
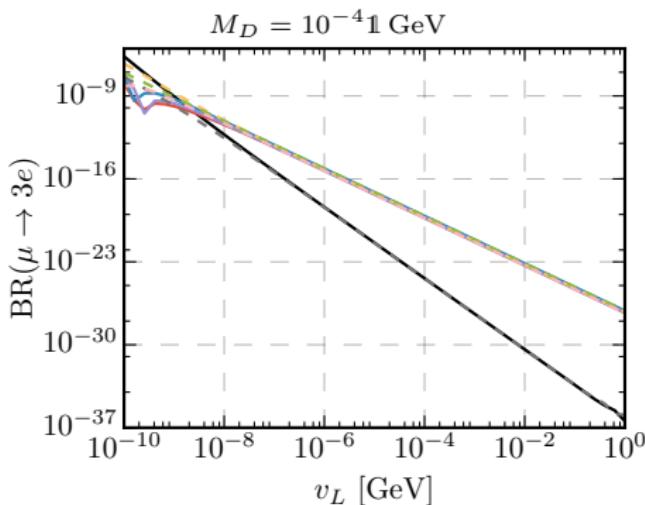


Scalar penguin



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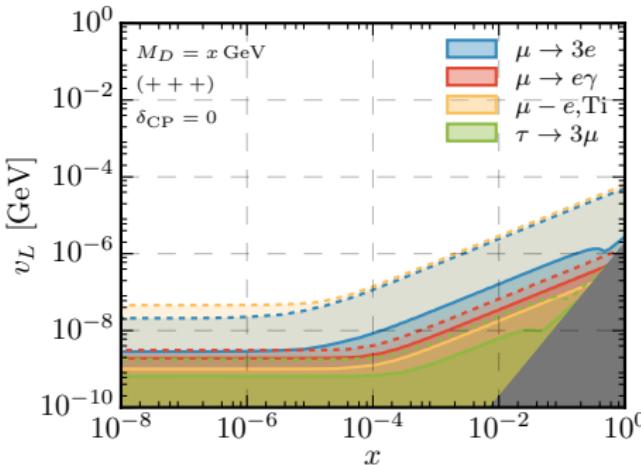
TRIPLET-YUKAWA SIGN-CHOICE



- Sufficient to consider only two choices

CURRENT AND FUTURE BOUNDS

DIAGONAL M_D



LFV Process	Present Bound	Future Sensitivity
$\mu \rightarrow e\gamma$	4.2×10^{-13}	6×10^{-14}
$\mu \rightarrow 3e$	1.0×10^{-12}	$\sim 3 \times 10^{-16}$
$\tau \rightarrow 3\mu$	2.1×10^{-8}	$\sim 4 \times 10^{-10}$
$\mu^- \rightarrow e^-, \text{Ti}$	4.3×10^{-12}	$\sim 10^{-18}$

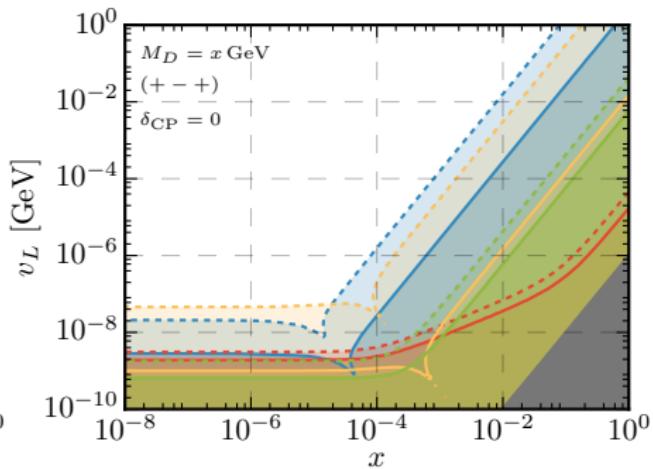
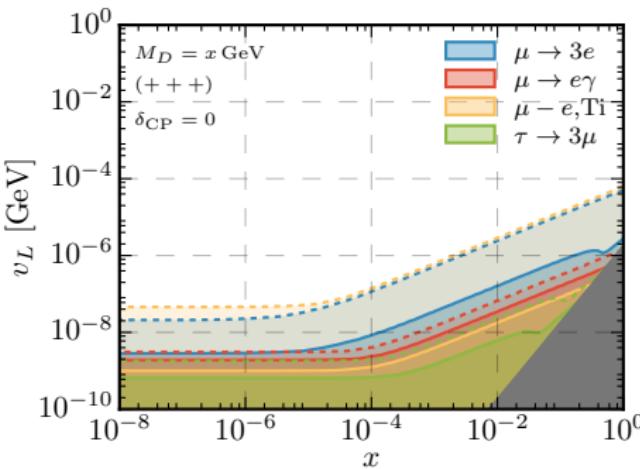
- Scalar tree-level contributions dominate
- Off-diagonal Y_Δ vanish at LO

$$D_B = \frac{m_\nu^{\text{light}}}{M_D} \ll \mathcal{O}(\sqrt{\alpha})$$

$$\Rightarrow D_A \simeq \pm \alpha^{-1/2}$$

$$\hookrightarrow A^{(++)} = V D_A V^T = D_A$$

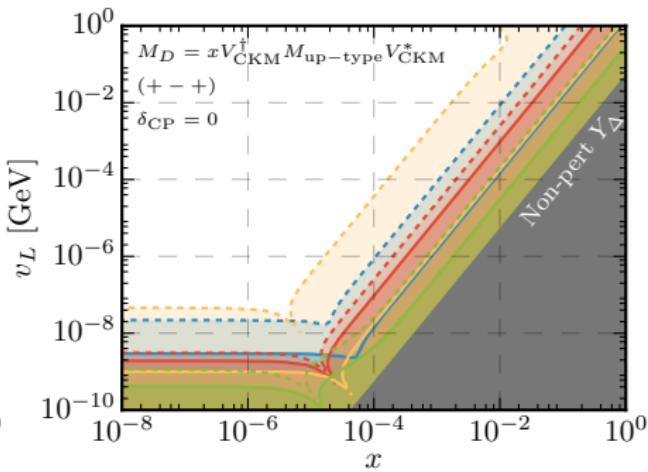
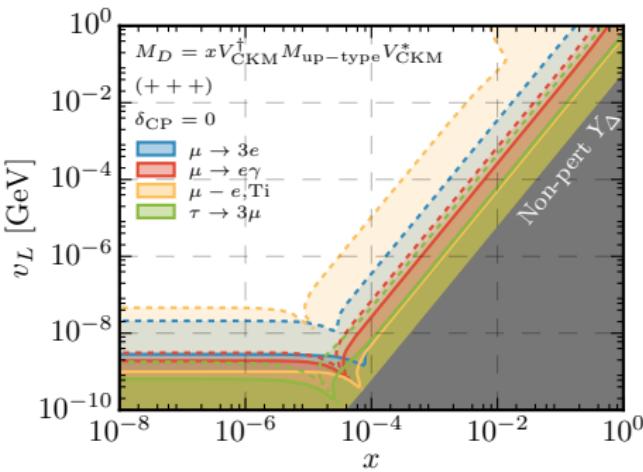
DIAGONAL M_D



- Scalar tree-level contributions again dominate
- Off-diagonal Y_Δ do not vanish at LO

$$A^{(+-)} = V \begin{pmatrix} \alpha^{-1/2} & 0 \\ 0 & -\alpha^{-1/2} \end{pmatrix} V^T = \alpha^{-1/2} \begin{pmatrix} \cos 2\theta & \sin 2\theta \\ \sin 2\theta & -\cos 2\theta \end{pmatrix} \neq D_A$$

$$M_D = V_{\text{CKM}}^\dagger M_{\text{up-type}} V_{\text{CKM}}^*$$



- Additional flavour violation from CKM elements
 - Large hierarchy in Y_Δ
 - Interplay between γ -penguins and tree-level scalar contributions
- Example $\mu \rightarrow 3e$:

$$\mathcal{M}_{\text{tree-level}} \propto Y_\Delta^{(2,1)} Y_\Delta^{(1,1)}$$

$$\mathcal{M}_{\gamma\text{-peng}} \propto \frac{1}{16\pi^2} \sum_i Y_\Delta^{(2,i)} Y_\Delta^{(i,1)}$$

CONCLUSIONS

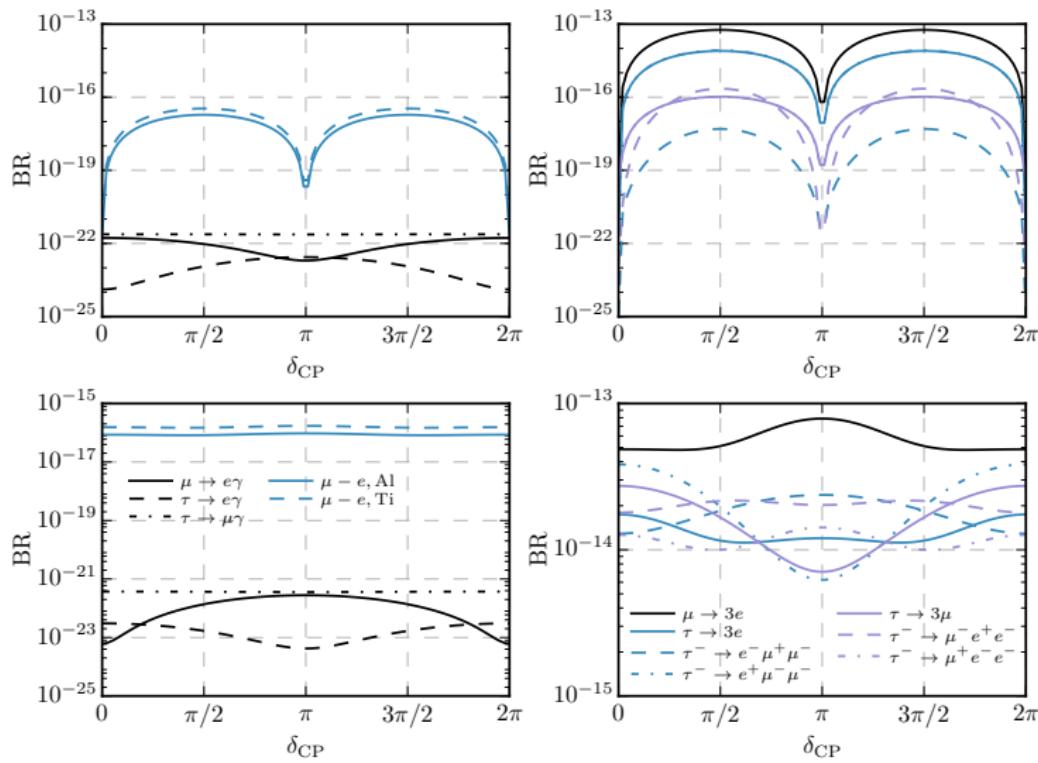
- ① Exploited discrete left-right symmetries to simplify neutrino sector
- ② Examined current constraints and prospects of probing low-scale left-right symmetries using LFV observables

Also investigated:

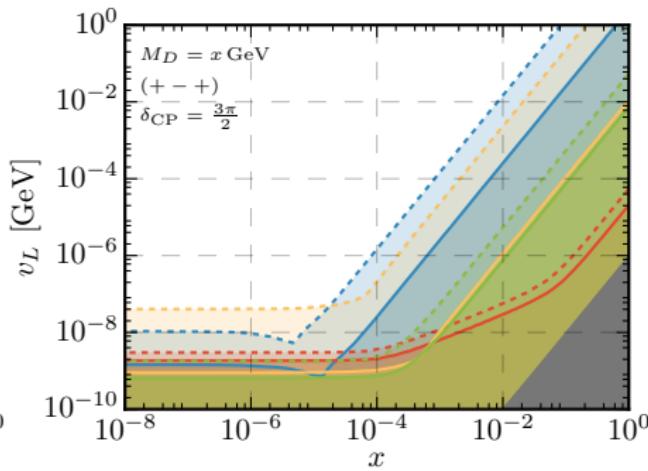
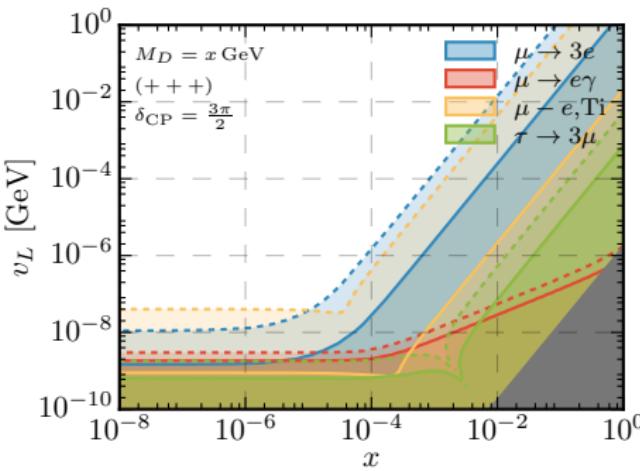
- Varying scalar masses
- Neutrino hierarchy & mass scale
- Varying δ_{CP}

BACKUP SLIDES

EFFECT OF δ_{CP}



EFFECT OF δ_{CP}



- For both sign choices off-diagonal Y_Δ **do not** vanish at LO

$$A^{(+-)} = V^* D_A V^\dagger \neq D_A$$

SCALAR SECTOR

Field content

$$\Phi = \begin{pmatrix} \phi_1^0 & \phi_1^+ \\ \phi_2^- & \phi_2^0 \end{pmatrix} \quad \Delta_L = \begin{pmatrix} \frac{\delta_L^+}{\sqrt{2}} & \delta_L^{++} \\ \delta_L^0 & -\frac{\delta_L^+}{\sqrt{2}} \end{pmatrix} \quad \Delta_R = \begin{pmatrix} \frac{\delta_R^+}{\sqrt{2}} & \delta_R^{++} \\ \delta_R^0 & -\frac{\delta_R^+}{\sqrt{2}} \end{pmatrix}$$

with VEVs

$$\begin{aligned} \phi_1^0 &= \frac{1}{\sqrt{2}} (v_1 + \sigma_1 + i\varphi_1) & \delta_L^0 &= \frac{1}{\sqrt{2}} (v_L + \sigma_L + i\varphi_L) \\ \phi_2^0 &= \frac{1}{\sqrt{2}} (v_2 + \sigma_2 + i\varphi_2) & \delta_R^0 &= \frac{1}{\sqrt{2}} (v_R + \sigma_R + i\varphi_R) \end{aligned}$$

SCALAR MASSES

Bi-doublet/SM-like masses

$$\begin{aligned} m_h^2 &\simeq 2\lambda_1 v^2 - \frac{8\lambda_4^2 v^4}{\alpha_3 v_R^2} & m_H^2 &\simeq 2(2\lambda_2 + \lambda_3)v^2 + \frac{\alpha_3}{2}v_R^2 \\ m_A^2 &\simeq 2\alpha_3 v_R^2 + 2(\lambda_3 - 2\lambda_2)v^2 & m_{H^\pm}^2 &\simeq \frac{1}{4}\alpha_3(v^2 + 2v_R^2) \end{aligned}$$

Left/right triplet masses

$$\begin{aligned} m_{H_L}^2 &\simeq \frac{1}{2}(\rho_3 - 2\rho_1)v_R^2 & m_{H_R}^2 &\simeq 2\rho_1 v_R^2 \\ m_{A_L}^2 &\simeq \frac{1}{2}(\rho_3 - 2\rho_1)v_R^2 & m_{H_L^\pm}^2 &\simeq \frac{1}{2}v_R^2(\rho_3 - 2\rho_1), \\ m_{H_1^{\pm\pm}}^2 &\simeq 2\rho_2 v_R^2 + \frac{1}{2}\alpha_3 v^2 & m_{H_2^{\pm\pm}}^2 &\simeq \frac{1}{2}\left((\rho_3 - 2\rho_1)v_R^2 + \alpha_3 v^2\right) \end{aligned}$$

LFV BOUNDS & FUTURE SENSITIVITY

LFV Process	Present Bound	Future Sensitivity
$\mu \rightarrow e\gamma$	4.2×10^{-13}	6×10^{-14}
$\tau \rightarrow e\gamma$	3.3×10^{-8}	$\sim 3 \times 10^{-9}$
$\tau \rightarrow \mu\gamma$	4.4×10^{-8}	$\sim 10^{-9}$
$\mu \rightarrow eee$	1.0×10^{-12}	$\sim 3 \times 10^{-16}$
$\tau \rightarrow eee$	2.7×10^{-8}	$\sim 5 \times 10^{-10}$
$\tau \rightarrow \mu\mu\mu$	2.1×10^{-8}	$\sim 4 \times 10^{-10}$
$\tau^- \rightarrow e^-\mu^+\mu^-$	2.7×10^{-8}	$\sim 5 \times 10^{-10}$
$\tau^- \rightarrow \mu^-e^+e^-$	1.8×10^{-8}	$\sim 3 \times 10^{-10}$
$\tau^- \rightarrow \mu^+e^-e^-$	1.5×10^{-8}	$\sim 3 \times 10^{-10}$
$\tau^- \rightarrow e^+\mu^-\mu^-$	1.7×10^{-8}	$\sim 3 \times 10^{-10}$
$\mu^- \rightarrow e^-, \text{Ti}$	4.3×10^{-12}	$\sim 10^{-18}$
$\mu^- \rightarrow e^-, \text{Au}$	7×10^{-13}	-
$\mu^- \rightarrow e^-, \text{Al}$	-	$10^{-16} - 3 \times 10^{-17}$
$\mu^- \rightarrow e^-, \text{SiC}$	-	10^{-14}

MODEL PARAMETERS AND MASS SPECTRUM

Model Parameters			
λ_1	0.13	v_L	$10^{-10} \dots 1 \text{ GeV}$
λ_2	1.0	v_R	20 TeV
λ_3	1.0	$\tan \beta$	10^{-4}
λ_4	0	α_1	0
ρ_1	3.2×10^{-4}	α_2	0
ρ_2	2.5×10^{-4}	α_3	2.0
ρ_3	1.8×10^{-3}	β_1	0
ρ_4	0	β_2	3.83×10^{-4}
μ_1^2	$7.87 \times 10^3 \text{ GeV}^2$	β_3	0
μ_2^2	$-2.00 \times 10^4 \text{ GeV}^2$	μ_3^2	$1.28 \times 10^5 \text{ GeV}^2$
Resulting Mass Spectrum			
m_h	125.5 GeV	m_H	20 TeV
m_A	20 TeV	m_{H^\pm}	20 TeV
m_{H_L}	482 GeV	m_{H_R}	506 GeV
m_{A_L}	482 GeV	$m_{H_L^\pm}$	512 GeV
$m_{H_1^{\pm\pm}}$	511 GeV	$m_{H_2^{\pm\pm}}$	541 GeV
M_{W_R}	9.37 TeV	M_{Z_R}	15.7 TeV